# A Simplified Clear Sky Model for Direct and Diffuse Insolation on Horizontal Surfaces

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# **Solar Energy Research Institute**

A Division of Midwest Research Institute

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A SIMPLIFIED CLEAR SKY MODEL FOR DIRECT AND DIFFUSE INSOLATION ON HORIZONTAL SURFACES

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#### PREFACE

The work documented here was performed by the SERI Renewable Resource Assessment Branch for the U.S. Department of Energy under Task No. 1093.00. The report compares several simple global horizontal insolation models with several rigorous radiative transfer models and describes an improved, simple, global insolation model. We would like to thank J. V. Dave of IBM for providing data sets from his Spherical Harmonics code.

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# **NOMENCLATURE**

a <sub>i</sub>	Absorptance parameters defined by Hoyt
a <sub>co</sub>	Carbon dioxide absorptance
aox	Oxygen absorptance
a <sub>w</sub>	Water vapor absorptance
Ba	Ratio of the forward-scattered irradiance to the total scattered irradiance due to aerosols
Fzj	A parameter used by Watt in air mass calculations
Ias	Solar irradiance on a horizontal surface from atmospheric scattering ( $\rm W/m^2$ )
$I_d$	Direct solar irradiance on a horizontal surface $(W/m^2)$
$\mathbf{I}_{\mathbf{G}}$	Solar irradiance on a horizontal surface from multiple reflections between the ground and sky $(W/m^2)$
Is	Solar irradiance on a horizontal surface from scattered light (I $_{\mbox{\scriptsize as}}$ + I $_{\mbox{\scriptsize G}})$
IT	Total (global) solar irradiance on a horizontal surface $(W/m^2)$
I <sub>o</sub>	Extraterrestrial solar irradiance (1353 $W/m^2$ )
K	Aerosol transmittance for Davies and Hay
к <sub>1</sub>	Constant used in Bird model associated with aerosol absorptance
М	Air mass
M '	Pressure-corrected air mass
$^{\rm M}{ ilde{ ilde{ imes}}}$	Watt's path length modifier (similar to air mass)
P	Surface pressure (millibars)
rg	Ground albedo
$r_s$	Sky, or atmospheric, albedo
T	Surface temperature (K)
$T_{A}$	Transmittance of aerosol absorptance and scattering
T <sub>AA</sub>	Transmittance of aerosol absorptance



TAS	Transmittance of aerosol scattering
Tas	Transmittance of dry air absorptance and scattering for Watt
$\mathtt{T}_{\mathbf{L}}$	Transmittance of lower layer aerosol for Watt
T <sub>M</sub>	Global transmittance of all molecular effects except water vapor for Atwater
T <sub>Md</sub>	Direct transmittance of all molecular effects except water vapor for Atwater
To	Transmittance of ozone absorptance
$T_{\mathbf{R}}$	Transmittance of Rayleigh scattering
$T_{\overline{U}}$	Transmittance of upper layer aerosol for Watt
T <sub>UM</sub>	Transmittance of absorptance of uniformly mixed gases (carbon dioxide and oxygen)
T <sub>w</sub>	Transmittance of water vapor absorptance $(1 - a_w)$
T <sub>ws</sub>	Transmittance of water vapor scattering
Uo	Amount of ozone in a vertical column from surface (cm)
$\mathbf{u}_{\mathbf{w}}$	Amount of precipitable water in a vertical column from surface (cm)
x <sub>o</sub>	Total amount of ozone in a slanted path (cm)
$X_{W}$	Total amount of precipitable water in a slanted path (cm)
Z	Angle between a line to the sun and the local zenith (zenith angle in degrees)
α <sub>cs</sub>	An attenuation multiplier used by Watt
$ au_{ extbf{A}}$	Broadband aerosol optical depth from surface in a vertical path (broadband turbidity)
τ0.5	Aerosol optical depth from surface in a vertical path at $0.5\mu\text{m}$ wavelength
τ0.38	Aerosol optical depth from surface in a vertical path at 0.38- $\mu\text{m}$ wavelength
$\tau_L$	Lower layer aerosol optical depth in a vertical path for Watt
$\tau_{\mathbf{u}}$	Upper layer aerosol optical depth in a vertical path for Watt
W <sub>O</sub>	Single scattering albedothe fraction of the light lost from an incident pencil of radiation that is due to scattering



#### SECTION 1.0

#### INTRODUCTION

To properly design a solar energy system for a location lacking an insolation data base, insolation models are required. This need for insolation models has been recognized for many years. One early and widely used insolation model was published in 1940 by Moon [1]. This model is still used today in its original or modified forms [2,3].

Insolation models have proliferated to the point where it is difficult for a solar user to decide which model to adopt. The purpose of this paper is to provide a detailed comparison of several simple, broadband insolation models that are currently in use and, based on this comparison and comparisons with more rigorous radiative transfer models, to formulate a simple clear sky model for direct and diffuse insolation. This type of comparison should be helpful in evaluating the relative accuracy of various models and give direction for formulating a model that uses the best of each existing model. The criteria used for evaluating and formulating models have been simplicity, accuracy, and the ability to use readily available meteorological data.

It should be noted that the use of the word "rigorous" does not necessarily mean that the "rigorous" results truly represent reality. Even though these "rigorous" codes are very detailed in the methods used to solve the radiative transfer problem, the representativeness of the results depends upon how well the atmospheric model, the measured atmospheric parameters, the mathematical methods, and other assumptions made in the codes relate to a real situation. However, in the absence of well-documented data, this approach of using "rigorous" codes as a basis of comparison was used.



#### SECTION 2.0

#### DESCRIPTION OF MODELS

A brief description of the several models that have been compared is presented here. Further details of this comparison can be found in Ref. 4 and in the original publications of each author. Most of the models include the effect of clouds, but this aspect of the models is not included here. Comparisons of cloud-cover global insolation models will follow (in a subsequent report) the clear sky model comparisons reported here. Such comparisons have also been performed by Davies and Hay [5].

#### 2.1 ATWATER AND BALL MODEL

Direct and global insolation models were published by Atwater and Ball [6,7]. The direct insolation model was taken from Kastrov as discussed by Kondratyev [8]. The equations of transfer and the transmission functions for this insolation model are given in Table 2-1. (The symbols in Tables 2-1 through 2-6 are defined in the Nomenclature.)

Table 2-1. EQUATIONS FOR TOTAL DOWNWARD IRRADIANCE FOR ATWATER AND BALL MODEL

# Basic Equations

$$I_d = I_o (\cos Z) (T_{Md} - a_w) T_A$$

$$I_{T} = I_{o} (\cos Z) (T_{M} - a_{w})T_{A}/(1 - r_{g}r_{s})$$

# Transmission Functions

$$T_{Md} = 1.041 - 0.16 [M(949 \times 10^{-6} P + 0.051)]^{0.5}$$

$$T_{M} = 1.021 - 0.0824 [M(949 \times 10^{-6} P + 0.051)]^{0.5}$$

$$a_w = 0.077 (U_w M)^{0.3}$$

$$T_A = \exp(-\tau_A M')$$

$$M = 35/[(1224 \cos^2 Z) + 1]^{0.5}$$

M'' = PM/1013

The form of the equation for water vapor absorption was published by McDonald [9]. The value of  $r_s = 0.0685$  for a molecular atmosphere, as reported by Lacis and Hansen [10], was used with this model. Atwater and Ball



used MIE calculations to obtain  $\tau_A$ , which is much too rigorous for a simple model. Therefore, a value of  $\tau_A$  that will be described later in the Bird model was used here.

#### 2.2 DAVIES AND HAY MODEL

A model for solar insolation (direct and diffuse) was published by Davies and Hay [5]. The equations used in this model were partially the result of comparing several existing models, and they are presented in Table 2-2.

Table 2-2. EQUATIONS FOR TOTAL DOWNWARD IRRADIANCE FOR DAVIES AND HAY MODEL

# Basic Equations

$$I_{d} = I_{o} (\cos Z) (T_{o}T_{R} - a_{w})T_{A}$$

$$I_{as} = I_{o} (\cos Z) [T_{o}(1 - T_{R})T_{A} (0.5) + (T_{o}T_{R} - a_{w})(1 - T_{A})W_{o}B_{a}]$$

$$I_{G} = r_{g}r_{s}(I_{d} + I_{as})/(1 - r_{g}r_{s})$$

$$I_{T} = I_{d} + I_{as} + I_{G}$$

# Transmission Functions

$$T_{o} = 1 - 0.02118X_{o}/(1 + 0.042X_{o} + 0.000323X_{o}^{2})$$

$$-1.082X_{o}/(1 + 138.6X_{o})^{0.805} - 0.0658X_{o}/[1 + (103.6X_{o})^{3}]$$

$$X_{o} = U_{o}M$$

$$a_{w} = 2.9X_{w}/[(1 + 141.5X_{w})^{0.635} + 5.925X_{w}]$$

$$X_{w} = U_{w}M$$

$$T_{A} = K^{M}$$

$$r_{S} = 0.0685 + (1 - B_{a})(1 - T_{A})W_{o}$$

The expressions for ozone transmittance,  $T_{\rm O}$ , and water vapor absorption,  $a_{\rm W}$ , were taken from Lacis and Hansen [10]. The transmission due to Rayleigh scattering,  $T_{\rm R}$ , was presented in tabular form, and so we used the Bird model expression for  $T_{\rm R}$  in this model. The value K = 0.91 was used for data generated here and is representative of aerosol conditions in southern Ontario.  $W_{\rm O}$  = 0.98 and  $B_{\rm a}$  = 0.85 were used here also.



#### 2.3 WATT MODEL

Another direct and diffuse insolation model has been constructed by Watt [3], based partially on the work of Moon [1]. The equations for this model are shown in Table 2-3.

Table 2-3. EQUATIONS FOR TOTAL DOWNWARD IRRADIANCE FOR WATT MODEL

# Basic Equations

$$I_{d} = I_{o}(\cos Z) T_{wa}T_{as}T_{o}T_{ws}T_{L}T_{U}$$

$$I_{s} = I_{o}[0.8 r_{s}(1 + r_{g}r_{s})(1 + \cos Z)^{0.5} + 0.5 \alpha_{cs}r_{g}r_{s} \cos Z + 0.5 r_{s} \cos Z]$$

$$I_{T} = I_{d} + I_{s}$$

# Transmission Functions

$$T_{w} = 0.93 - 0.033 \log (U_{w}M_{2})$$

$$T_{as} = 10^{-0.045[(P/P_{o}) M_{1}]^{0.7}}$$

$$T_{o} = 10^{-(0.0071 + 0.01 U_{o}M_{4})}$$

$$T_{ws} = 10^{-(0.0095 U_{w}M_{2})}$$

$$T_{L} = 10^{T}L^{M}2^{0.7}$$

$$T_{U} = 10^{-T}u^{M}3$$

$$\tau_{L} = 0.6 (\tau_{0.5} - 0.01 U_{w} - 0.03)$$

$$\alpha_{cs} = (0.93 - 0.033 \log U_{w}) 10^{-[0.006 P/1013 + 0.4 (T_{L} + T_{U})]}$$

$$r_{s} = \alpha_{cs} \left\{1 - 10^{-[0.003 P/1013 + 0.01 U_{w} + 0.4 (T_{L} + T_{U})]\right\}$$

$$M_{i} = \sec z \text{ for } z < 70^{\circ} (i = 1,2,3,4)$$

$$M_{i} = (h_{2}F_{z2} - h_{1}F_{z1})/(h_{2} - h_{1}) \text{ for } z > 70^{\circ} (i = 1,2,3,4)$$

$$F_{zj} = \left[(r/h_{j}) \cos z\right]^{2} + 2 r/h_{j} + 1\right]^{0.5} - (r/h_{j}) \cos z \quad (j = 1,2)$$

$$for i = 1, h_{1} = 0 \text{ km and } h_{2} = 30 \text{ km}$$

$$i = 2, h_{1} = 0 \text{ km and } h_{2} = 30 \text{ km}$$

$$i = 3, h_{1} = 15 \text{ km and } h_{2} = 25 \text{ km}$$

$$i = 4, h_{1} = 20 \text{ km and } h_{2} = 40 \text{ km}$$



The upper layer broadband turbidity,  $\tau_u$ , was not well defined by Watt. A value of  $\tau_u$  = 0.02 was used in the calculations performed here, which appears to be an average value for locations in the United States. The parameter r is the earth's radius (6.4 × 10 m).

## 2.4 HOYT MODEL

The equations used in the model by Hoyt [11] are shown in Table 2-4.

Table 2-4. EQUATIONS FOR TOTAL DOWNWARD IRRADIANCE FOR HOYT MODEL

$$I_{d} = I_{o} (\cos Z) \left( 1 - \sum_{i=1}^{5} a_{i} \right) T_{AS} T_{R}$$

$$I_{as} = I_{o} (\cos Z) \left( 1 - \sum_{i=1}^{5} a_{i} \right) [(1 - T_{R})0.5 + (1 - T_{AS})0.75]$$

$$I_{G} = (I_{d} + I_{as}) r_{g} \left( 1 - \sum_{i=1}^{5} a_{i} \right) [(1 - T_{R}')0.5 + (1 - T_{AS}')0.25]$$

$$I_{T} = I_{d} + I_{as} + I_{G}$$

## Transmission Functions

$$a_1 = a_w = 0.110 (0.75 U_w M + 6.31 \times 10^{-4})^{0.3} - 0.0121$$
 $a_2 = a_{co} = 0.00235 (126 M' + 0.0129)^{0.26} - 7.5 \times 10^{-4}$ 
 $a_3 = (1 - T_o) = 0.045 (U_o M + 8.34 \times 10^{-4})^{0.38} - 3.1 \times 10^{-3}$ 
 $a_4 = a_{ox} = 7.5 \times 10^{-3} (M')^{0.875}$ 
 $a_5 = 0.05 T_{AS}$ 
 $M' = MP/1013.25$ 

Hoyt obtained air mass values, M, from Bemporad's [12] tables. The expression for air mass of Kasten [13] was used here instead. The values of  $T_{\rm AS}$  and  $T_{\rm R}$  are calculated from tables furnished by Hoyt [11]. The  $a_{\rm i}$ ' values are calculated using air mass values of M' + 1.66 P/1013.25 in the  $a_{\rm i}$  expressions.



 $T_{AS}$ ' and  $T_{R}$ ' are evaluated for air mass values of 1.66 P/1013.25 in  $T_{AS}$  and  $T_{R}$ . The table values from which  $T_{AS}$  is calculated are limited so that large optical depths cannot be considered. Large optical depths can occur from high turbidity or from large zenith angles. In the data presented later for Z = 80°, an approximate value of  $T_{AS}$  was used to complete the plotted results.

## 2.5 LACIS AND HANSEN MODEL

The equations for the model developed by Lacis and Hansen [10] are shown in Table 2-5.

Table 2-5. EQUATIONS FOR TOTAL DOWNWARD IRRADIANCE FOR LACIS AND HANSEN MODEL

# Basic Equation

$$I_T = I_o (\cos Z) [(0.647 - r_s' - a_o)/(1 - 0.0685 r_g) + 0.353 - a_w]$$

## Transmission Functions

$$r_s' = 0.28/(1 + 6.43 \cos Z)$$

$$a_0 = 1 - T_0$$
 as shown in Table 2-2

$$a_{\rm w}$$
 = Shown in Table 2-2 with the following correction:

$$X_w = X_w (P/1013)^{0.75} (273/T)^{0.5}$$

#### 2.6 BIRD MODEL

A model has been constructed that is based on comparisons with the SOLTRAN 3 and SOLTRAN 4 [4] direct insolation models and the BRITE Monte Carlo global model [14]. Formalisms in the previous models that were considered to be optimum were adopted here. The equations for this model are shown in Table 2-6.

The atmospheric turbidity values,  $\tau_{A,0.38}$  and  $\tau_{A,0.5}$ , have been measured on a regular basis by the National Weather Service [15] at 0.38- and 0.5- $\mu$ m wavelengths, respectively. If one of the turbidity values is not available, its value can be entered as a zero in the expression for  $\tau_A$ . The expression for  $\tau_A$  is based on the Air Force Geophysics Laboratory (AFGL) rural aerosol model [16]. The expression used here for  $\tau_{AA}$  was found by fitting the expression to the results of the SOLTRAN 4 [4] code. The value of  $\kappa_1$  = 0.0933 for the rural aerosol was used here for all calculations. For the urban aerosol model, which contains more carbon, the value of  $\kappa_1$  = 0.385 was found to be appropriate. From a theoretical standpoint,  $\kappa_1$  should be nearly equal to  $1-\nu_0$ , where  $\nu_0$  is the single scattering albedo. The forward-scattering



## Table 2-6. EQUATIONS FOR TOTAL DOWNWARD IRRADIANCE FOR THE BIRD MODEL

## Basic Equations

$$I_{d} = I_{o} (\cos Z) (0.9662) T_{R} T_{o} T_{UM} T_{w} T_{A}$$

$$I_{as} = I_{o} (\cos Z) (0.79) T_{o} T_{w} T_{UM} T_{AA}$$

$$[0.5 (1 - T_{R}) + B_{a} (1 - T_{AS})] / [1 - M + (M)^{1.02}]$$

$$I_{T} = (I_{d} + I_{as}) / (1 - r_{g} r_{s})$$

# Transmission Equations

$$T_{R} = \exp \left\{-0.0903 \text{ (M')}^{0.84}[1 + \text{M'} - (\text{M'})^{1.01}]\right\}$$

$$T_{O} = 1 - 0.1611 \text{ X}_{O} (1 + 139.48 \text{ X}_{O})^{-0.3035} - 0.002715 \text{ X}_{O} (1 + 0.044 \text{ X}_{O} + 0.0003 \text{ X}_{O}^{2})^{-1}$$

$$X_{O} = U_{O}M$$

$$T_{UM} = \exp \left[-0.0127 \text{ (M')}^{0.26}\right]$$

$$T_{W} = 1 - 2.4959 \text{ X}_{W} \left[(1 + 79.034 \text{ X}_{W})^{0.6828} + 6.385 \text{ X}_{W}\right]^{-1}$$

$$X_{W} = U_{W}M$$

$$T_{A} = \exp \left[-\tau_{A}^{0.873} (1 + \tau_{A} - \tau_{A}^{0.7088}) \text{ M}^{0.9108}\right]$$

$$\tau_{A} = 0.2758 \tau_{A,0.38} + 0.35 \tau_{A,0.5}$$

$$T_{AA} = 1 - K_{1}(1 - M + M^{1.06})(1 - T_{A})$$

$$T_{AS} = T_{A}/T_{AA}$$

$$r_{S} = 0.0685 + (1 - B_{a})(1.0 - T_{aS})$$

$$M = \left[\cos z + 0.15(93.885 - z)^{-1.25}\right]^{-1}$$

$$M' = MP/1013$$

ratio,  $B_a$ , is related through MIE theory to a parameter  $(\cos \theta)$ , called the asymmetry factor, by

$$B_a = 0.5(1 + (\cos \theta))$$
.



The asymmetry factor is the mean of the cosine of the scattering angle,  $\theta$ , with the angular intensity as the weighting function. The extreme values of  $B_a$  are

Table 2-7 contains values of the asymmetry factor at various wavelengths for the rural aerosol model and the haze L aerosol model used by Dave [17]. Two differences in these aerosol models are: (1) the rural aerosol model is bimodal, whereas Dave's model has a single mode; and (2) the rural aerosol model varies the complex index of refraction with wavelength, and the Dave model holds it constant. The values of the asymmetry factors for the two models are in reasonably good agreement, and our calculations indicate that the Bird model is relatively insensitive to small changes in this parameter. A value of  $B_{\rm a} = 0.82$  was used for the rural aerosol, and  $B_{\rm a} = 0.86$  for Dave's aerosol in calculations shown later.

Table 2-7. VALUES OF THE ASYMMETRY FACTOR FOR THE RURAL AND THE DAVE HAZE L AEROSOL

Rur	al	Dave Haze L				
λ	<cos θ=""></cos>	λ	<cos θ=""></cos>			
0.325	0.66	••••				
0.35	0.66	0.35	0.73			
0.4	0.65		<b></b> '			
0.5	0.64	0.455	0.72			
0.63	0.64	0.635	0.71			
0.7525	0.63	0.7525	0.71			
0.86	0.63		<del></del> _			
0.9935	0.63	0.994	0.70			
1.235	0.64	1.235	0.69			
1.497	0.65	1.61	0.67			
1.8	0.68	2.1	0.63			
2.198	0.71	2.198	0.62			

It is suggested that values of  $B_a = 0.84$  and  $K_1 = 0.1$  be used with this model unless good information on the aerosol is available. All other data required by the model comes from meteorological measurements near the site of interest.

Ozone measurements are sometimes difficult to obtain. Since ozone has a minor effect on broadband solar insolation, it is suggested that the method of Van Heuklon [18] could be used in lieu of real data.



The expression used for air mass in the Bird model comes from Kasten [13] and was used for all of the calculations reported here except for the Atwater and Ball model and the Watt model.

For the convenience of the reader, Table 2-8 itemizes the input parameters required for each of the simple models.

Table 2-8. INPUT PARAMETERS REQUIRED FOR SIMPLE MODELS

Model	Input
Atwater and Ball	Solar constant, zenith angle, surface pressure, ground albedo, precipitable water vapor, total ozone, broadband turbidity
Davies and Hay	Solar constant, zenith angle, surface pressure, ground albedo, precipitable water vapor, total ozone, aerosol single scattering albedo (0.98 suggested), aerosol forward scattering ratio (0.85 suggested), broadband aerosol transmittance
Watt	Solar constant, zenith angle, surface pressure, ground albedo, precipitable water vapor, total ozone, turbidity at 0.5-µm wavelength, upper layer turbidity
Hoyt	Solar constant, zenith angle, surface pressure, ground albedo, precipitable water vapor, total ozone, turbidity at one wavelength
Lacis and Hansen	Solar constant, zenith angle, surface pressure, surface temperature, ground albedo, precipitable water vapor, total ozone
Bird	Solar constant, zenith angle, surface pressure, ground albedo, precipitable water vapor, total ozone, turbidity at 0.5- and/or 0.38-µm wavelength, aerosol forward scattering ratio (0.84 recommended)

#### 2.7 RIGOROUS CODES

Three rigorous codes have been used in this study as a basis for formulating the Bird model. One code is for direct normal irradiance and is called SOLTRAN 4 [4]. Two other codes, which include both the direct normal and the diffuse irradiance, are the BRITE [14] Monte Carlo code and the Dave [17] Spherical Harmonics code. In each of these codes, a multilayered atmospheric model can be constructed, where important atmospheric parameters are defined



at each layer. In this way, a fairly detailed atmosphere can be constructed that closely resembles the real atmosphere at a given time and specific location. Each code then uses its own technique to solve the radiative transfer problem.

These rigorous codes calculate the irradiance at a specified altitude, sun angle, and for an atmospheric model at discrete wavelengths. For comparison with the simple broadband models described previously, the spectral irradiance from the rigorous codes has to be integrated over wavelength.

The SOLTRAN 3 code, an earlier version of SOLTRAN 4, was used to formulate most of the transmittance functions found in the Bird model. This was accomplished by performing a least-square fit of each transmittance function to transmittance data from SOLTRAN 3 for each atmospheric constituent. Further details of this operation can be found in Ref. 4. SOLTRAN 3 was used because it was the only version available to us when this portion of the work was performed. The only difference in the two models that would be apparent in the results would be caused by a slight difference in the "continental" and "rural" aerosol models that are used in the two codes.

A comparison was made between data from the BRITE code and several of the simple clear sky global models. Based on this comparison and the author's judgment of the best expressions used in the simple models, a model for the diffuse irradiance was formulated. This simple model of the diffuse irradiance was then fine-tuned to provide good agreement with the BRITE code as well as results from the Dave Spherical Harmonics code.

It is appropriate to comment here that there are problems with the expressions used for the diffuse irradiance model. The general formalism for the diffuse transfer equation of some of the simple models was adopted even though it may not be as acceptable, based on the physics of the problem, as one would like. For example, a cosine of the solar zenith angle is included in the diffuse transfer equation. This implies that all of the diffuse radiation behaves just like the direct normal component. The cosine is used to calculate the irradiance falling on a horizontal surface. However, it is well known that the diffuse irradiance is much more complex than this. An example of a more rigorous but fairly simple formalism for the diffuse irradiance is found in Ref. 19, in which the diffuse radiation is divided into three components: an isotropic term, a term resulting from horizon brightening, and The circumsolar term is the only one that behaves very a circumsolar term. much like the direct normal radiation. For tilted surfaces, a ground reflection should be added to this diffuse model. Another problem with this formalism is associated with using transmittance expressions for diffuse radiation that were derived for direct radiation.



#### SECTION 3.0

## MODEL COMPARISONS

Each of the simple models described in Section 2.0 was programmed on a computer to produce data for comparison. A comparison of the aerosol transmittance, the transmittance after molecular (Rayleigh) scattering, the water vapor  $({\rm H}_2{\rm O})$  transmittance, and the ozone  $({\rm O}_3)$  transmittance will be presented first. Then, a comparison between the direct, the diffuse sky, the diffuse sky/ground, and the global radiation for three different atmospheric models will be shown. Comparisons are made, where possible, between each of the simple model results as well as the results from the rigorous models.

### 3.1 TRANSMITTANCE COMPARISONS

To become oriented as to the relative importance of each atmospheric constituent in atmospheric transmittance, the broadband transmittance versus the secant of the solar zenith angle (approximate air mass) for each constituent was plotted in Fig. 3-1. This figure was generated with output from the SOLTRAN 3 code for a Midlatitude Summer (MLS) atmospheric model. Table 3-1 shows the amounts of  $\rm H_2O$  and  $\rm O_3$  from sea level to the top of the atmosphere in a vertical column for the two atmospheric models--MLS and USS (U.S. Standard)-used in this comparison.

Table 3-1. AMOUNTS OF H<sub>2</sub>O AND O<sub>3</sub> IN A VERTICAL COLUMN FOR THE MLS AND USS ATMOSPHERES

	H <sub>2</sub> 0 (cm)	03 (cm)
MLS	2.93	0.31
USS	1.42	0.34

An examination of Fig. 3-1 shows that  ${\rm CO_2}$  and  ${\rm O_2}$  are the least important attenuators, and this is why they are not included in some models. The next element exhibiting increased attenuation is  ${\rm O_3}$ , followed by  ${\rm H_2O}$ . Molecular scattering dominates the total molecular absorption at large zenith angles and has a greater effect than most individual molecular species at all zenith angles. The most significant attenuator at all zenith angles is the aerosol. The aerosol modeled here was the continental aerosol model [16] with a sea level meteorological range of 23 km. A 23-km meteorological range is considered to be a moderately clear atmosphere. This aerosol produces an optical depth of 0.27 (base e) at 0.5- $\mu$ m wavelength.



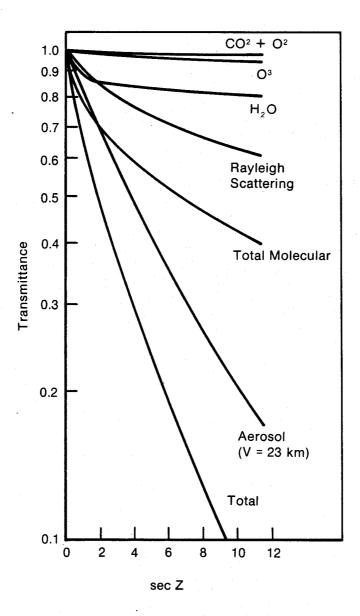


Figure 3-1. Transmittance versus Secant of Solar Zenith Angle for Midlatitude Summer Model



The data in Fig. 3-1 indicate the relative effect that the atmospheric constituents have on the direct irradiance. However, when global irradiance is being considered these results can be misleading. A large fraction of the radiation lost in the direct beam by molecular and aerosol scattering is regained in the diffuse component. As a result of this, changes in aerosol optical depth through the atmosphere have less effect on the total global irradiance than on the direct irradiance.

Figures 3-2 and 3-5 present transmittance data for the USS atmosphere from several of the models. It should be pointed out that the broadband turbidity expression from the Bird model was used in the Atwater and Ball model. A similar effort could have been made with the aerosol transmittance term in the Davies and Hay model to produce identical results. One of the strengths of the Bird model is that it is based entirely on algebraic expressions for transmittance calculations rather than tabulated data. This makes the use of the model considerably easier and provides more flexibility.

The comparison made here for one model atmosphere is not really indicative of the accuracy of each model. Since a wide range of values of turbidity,  $\rm H_2O$  amount, and  $\rm O_3$  amount are required for real atmospheric conditions, the model must be able to accommodate these changes. Additional comparisons are presented in Ref. 4 for a range of these parameters. As was stated earlier, the transmittance expressions in the Bird model were derived from comparisons with SOLTRAN 3 results, but the comparisons made here are with SOLTRAN 4 results. This should have an effect only on the aerosol transmittance shown in Fig. 3-2. The transmittance comparisons shown in Figs. 3-2 through 3-5 are rather self-explanatory, and so no further discussions are presented.

#### 3.2 IRRADIANCE COMPARISONS

The global solar irradiance has been divided into three components: the direct irradiance on a horizontal surface, the diffuse sky irradiance on a horizontal surface. The diffuse sky irradiance is the total diffuse radiation present when the ground has zero albedo (completely absorbing ground), and the diffuse ground/sky irradiance is that amount added to the total diffuse irradiance when the ground albedo is not zero.

Figures 3-6 through 3-9 present comparisons of the global irradiance at sea level in the USS atmosphere from all of the models as well as comparisons of the three components of the global irradiance. For this atmospheric model, the Bird, Hoyt, and Monte Carlo models produce very similar results. The model of Atwater and Ball significantly underestimates the diffuse sky irradiance, and the Watt model overestimates the diffuse sky irradiance. The results of the Davies and Hay model would have shown much better agreement with the Monte Carlo results if a more reasonable value of the aerosol transmittance had been used. The Bird model Rayleigh transmittance was used in the Davies and Hay model.

It is instructive to examine the relative magnitude of each component. For a solar zenith angle of zero (the sun directly overhead), the direct component provides approximately 81% of the total, the diffuse sky approximately 17%,



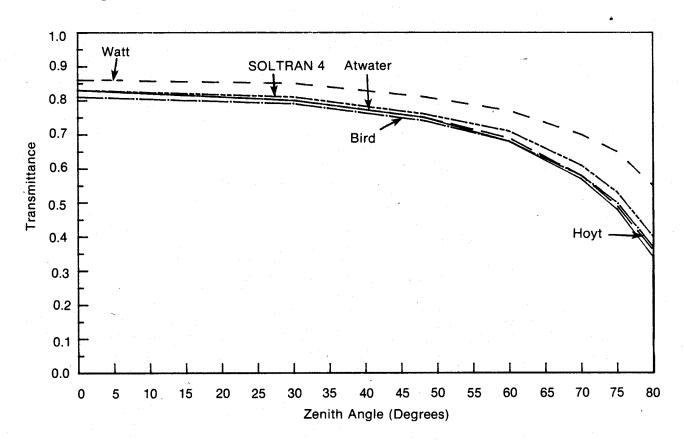


Figure 3-2. Aerosol Transmittance - USS Atmosphere (23-km-Visibility Aerosol)

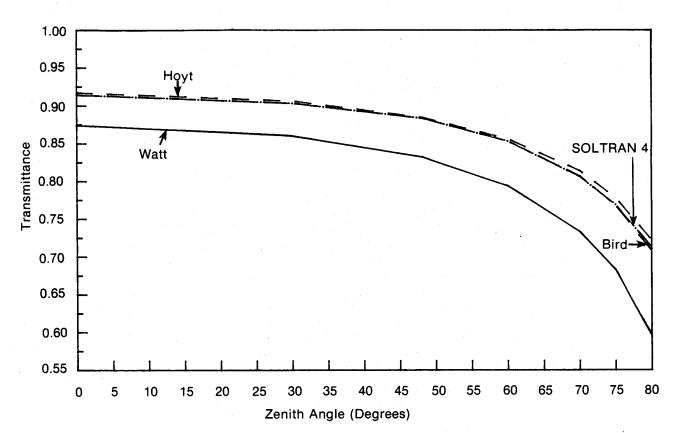


Figure 3-3. Rayleigh Transmittance



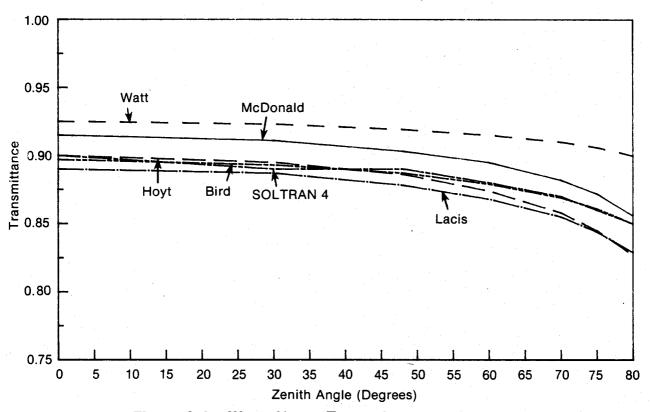


Figure 3-4. Water Vapor Transmittance – USS Atmosphere (1.42-cm H<sub>2</sub>O in Vertical Column)

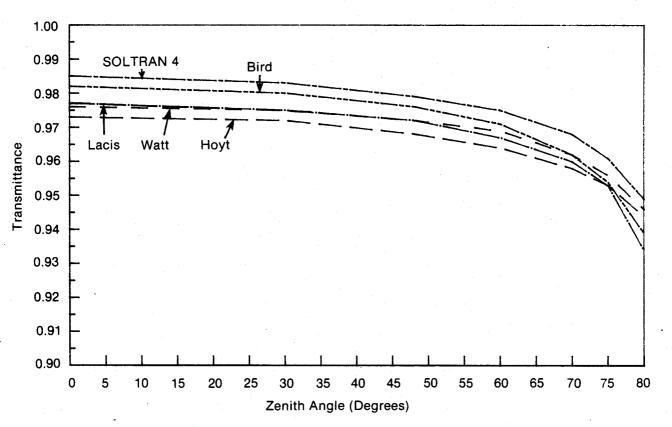


Figure 3-5. Ozone Transmittance — USS Atmosphere



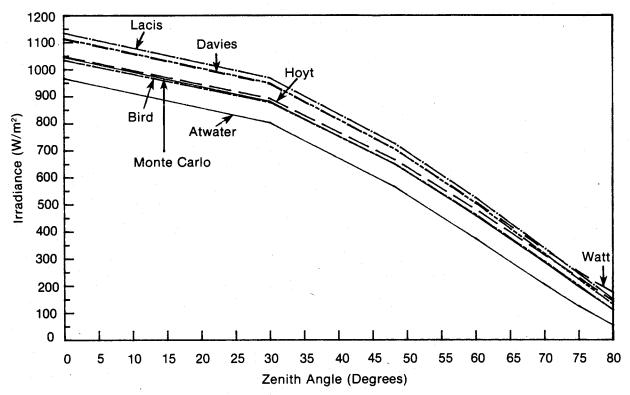


Figure 3-6. Total Insolation – USS Atmosphere (23-km-Visibility Rural Aerosol, Albedo = 0.2)

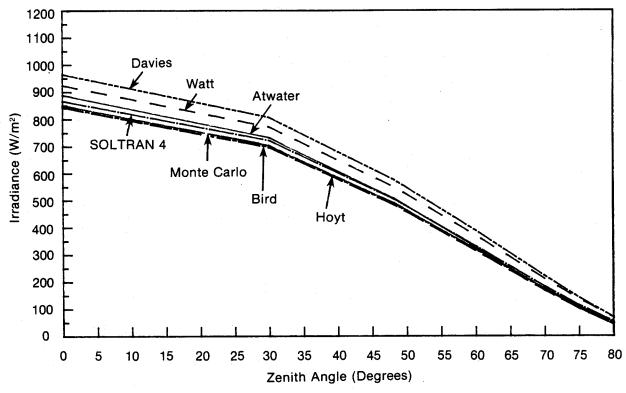


Figure 3-7. Direct Horizontal Insolation – USS Atmosphere (V = 23 km;  $U_W = 1.42$  cm;  $U_O = 0.34$  cm)



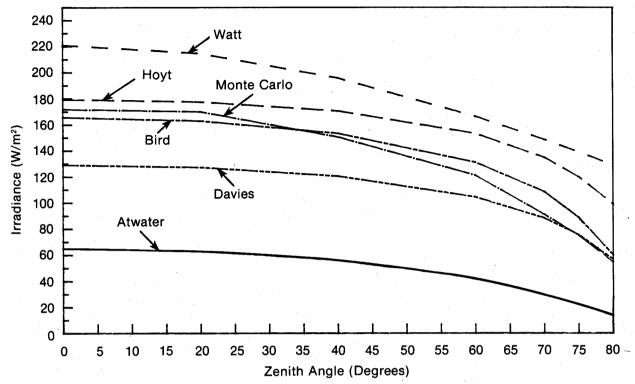


Figure 3-8. Diffuse Sky Insolation – USS Atmosphere (V = 23 km;  $U_w = 1.42$  cm;  $U_0 = 0.34$  cm)

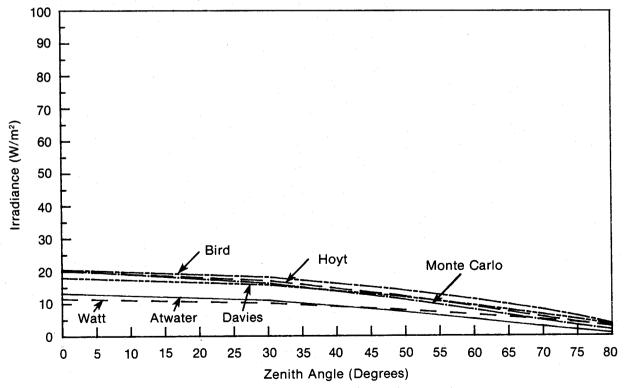


Figure 3-9. Diffuse Ground/Sky Insolation – USS Atmosphere (23-km-Visibility Rural Aerosol, Albedo = 0.2)



and the diffuse ground/sky approximately 2%. The calculations are for a turbidity of 0.27 at  $0.5-\mu m$  wavelength and a ground albedo of 0.2.

In Figures 3-10 through 3-13, a similar set of plots is presented for the MLS atmospheric model with an atmospheric turbidity of 0.27 at 0.5-µm wavelength and a ground albedo of 0.8. The larger ground albedo increases the diffuse ground/sky component by a factor of four. It is evident that the simple models begin to deviate from the Monte Carlo result for the diffuse ground/sky component. However, since this component is so small, the global result is still in close agreement. The relative agreement of the results from the different models is nearly the same as with the USS atmosphere. This is principally because the aerosol model is identical in both models.

Finally, a comparison is made of the models for the MLS atmosphere with the Haze L aerosol model of Dave [17]. The Dave atmosphere modeled here consists of 15 homogeneous layers instead of the 32 exponentially varying layers that were used in the previous MLS atmosphere. In addition, the Haze L aerosol model is significantly different from the rural aerosol model used previously. Not only are the particle size distributions and complex indices of refraction different, but most importantly the number density of the aerosol as a function of altitude is very different. The turbidity of this model is 0.0996 instead of the 0.266 used previously in the vicinity of 0.5- $\mu$ m wavelength. Calculations with SOLTRAN 4 show that a turbidity of 0.0996 in the rural aerosol model corresponds to a sea level visibility of nearly 250 km. This is an extremely clear atmosphere. Figures 3-14 through 3-17 illustrate the comparison results for this atmospheric model.

It is readily apparent from the results shown in Figs. 3-14 through 3-17 that the Atwater and Ball model is based on a very clear atmosphere, since it agrees much better with the Dave data. Similarly, the aerosol parameter, K=0.91, used in the Davies and Hay model is for a very clear atmosphere. The Lacis and Hansen model appears to be in slightly closer agreement with this clear atmosphere also, but it does not have provisions for changes in turbidity.

The clear sky diffuse irradiance of the Atwater and Ball model as shown in the figures presented here may be slightly lower than the model intended because of the way the calculations were performed. This model is really composed of two separate models: one for the direct irradiance and one for the global irradiance. The clear sky diffuse irradiance was obtained by running the model for a ground albedo of zero and then subtracting the direct horizontal from the global. If the direct horizontal irradiance is slightly high, as it appears to be, then the diffuse term would be slightly lower than expected. The real evaluation of this model should be made on the global horizontal irradiance.

It will be noted in Fig. 3-14 that there is a slight difference between the Monte Carlo global and the spherical harmonic global results of Dave (~3.6% at a zenith angle of 0). Figure 3-15 shows that a large fraction of this difference is in the direct component. Since the direct component of the Monte Carlo code is deterministic in nature rather than statistical, these differences are most likely due to differences in the molecular absorption coefficients used and the band absorption models used. Dave used an older set



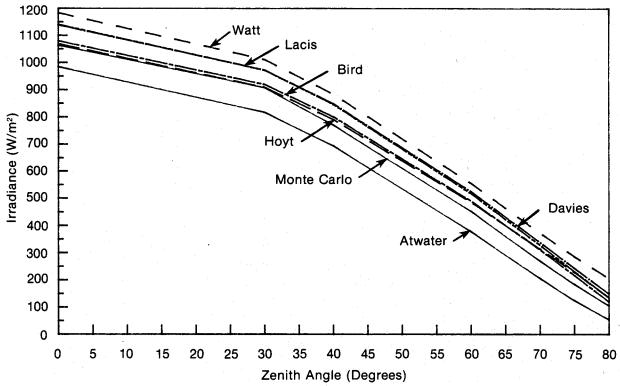


Figure 3-10. Total Insolation – MLS Atmosphere (23-km-Visibility Rural Aerosol, Albedo = 0.8)

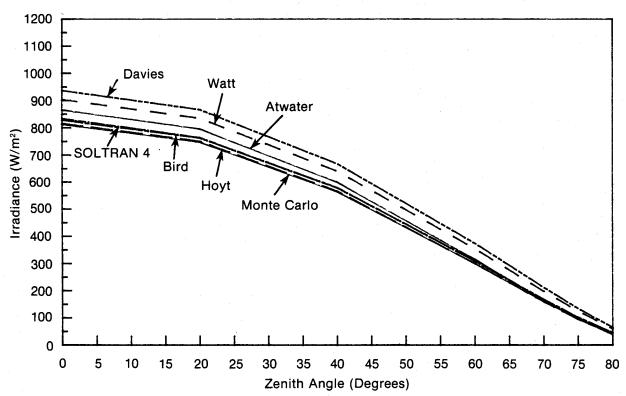


Figure 3-11. Direct Horizontal Insolation – MLS Atmosphere (V = 23 km, U<sub>W</sub> = 2.93 cm, U<sub>O</sub> = 0.31 cm)



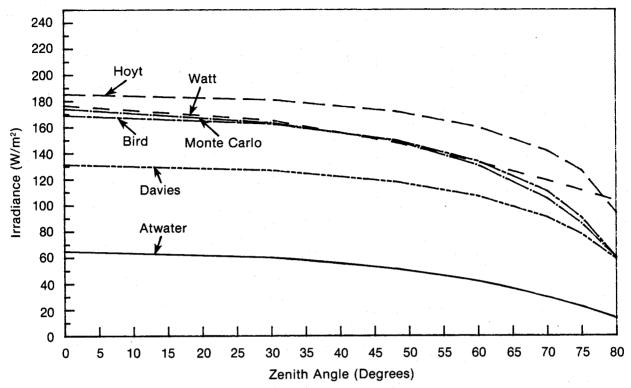


Figure 3-12. Diffuse Sky Insolation — MLS Atmosphere (V = 23 km,  $U_W = 2.93$  cm,  $U_O = 0.31$  cm)

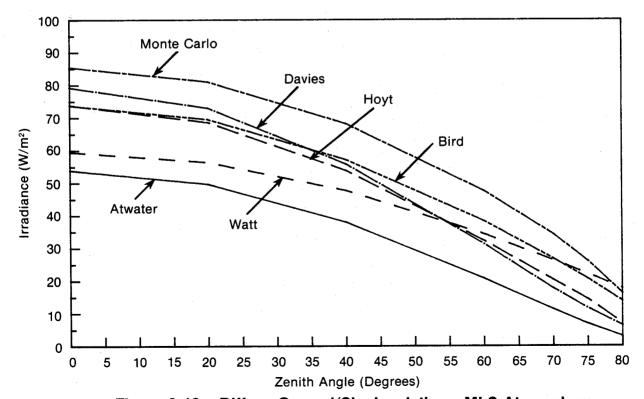


Figure 3-13. Diffuse Ground/Sky Insolation – MLS Atmosphere (23 - km-Visibility Rural Aerosol, Albedo = 0.8)



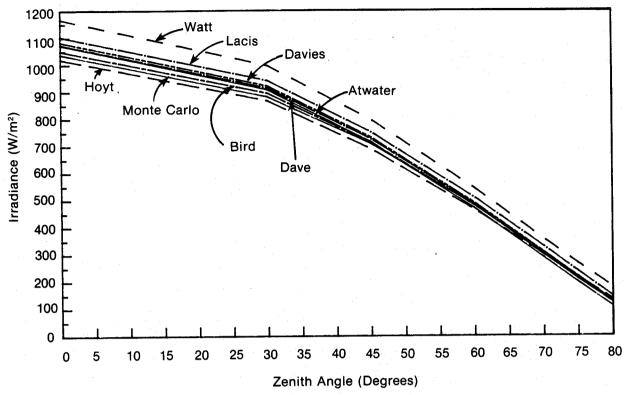


Figure 3-14. Total Insolation—Dave Model 3 (Albedo = 0.2)

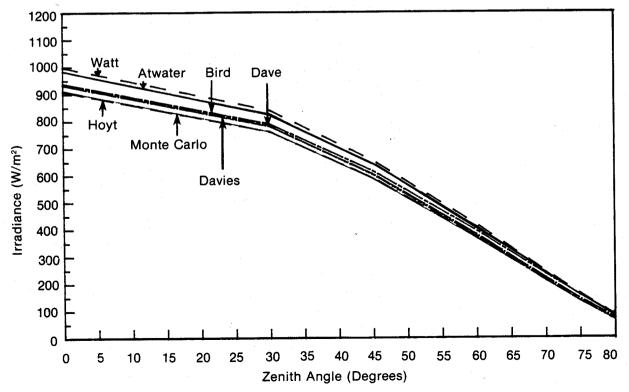


Figure 3-15. Direct Horizontal Insolation—Dave Model 3 (U<sub>W</sub> = 2.93 cm, U<sub>O</sub> = 0.31 cm)



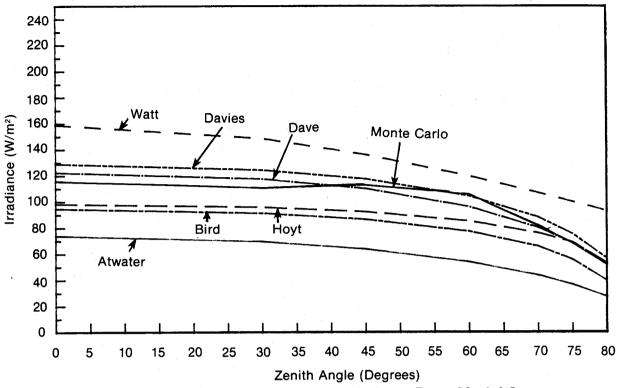


Figure 3-16. Diffuse Sky Insolation—Dave Model 3 (U<sub>W</sub> = 2.93 cm, U<sub>O</sub> = 0.31 cm)

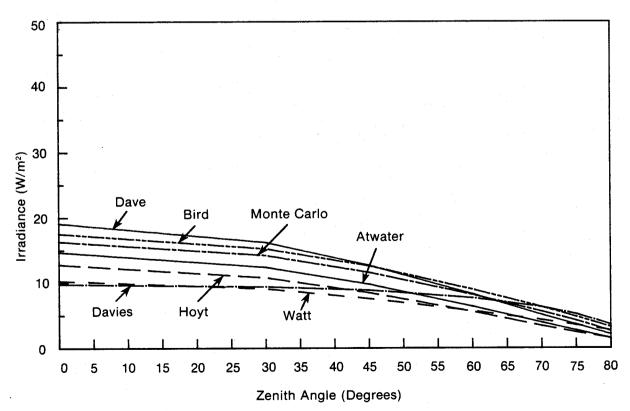


Figure 3-17. Diffuse Ground/Sky Insolation—Dave Model 3 (Albedo = 0.2)



of molecular absorption coefficients from AFGL than was used in the Monte Carlo code [20]. Figure 3-18 is a comparison of the spectral direct irradiance for these two codes. It is evident that there are some weak absorption bands present in the Monte Carlo code that are not present in the spherical harmonics code, and the shape of some of the bands is significantly different for the two codes. SOLTRAN 4 is based on the same absorption data that Dave used, and similar differences in the direct normal irradiance occur between SOLTRAN 4 and the Monte Carlo code. These differences are shown in Fig. 3-19, which has 31 more data points in the Monte Carlo results than in Fig. 3-18. This increase in the number of data points increases the apparent spectral resolution. The SOLTRAN 4 code provided approximately 600 data points in this figure. Our conclusion is that most of the differences in the results from the rigorous codes are due to differences in the molecular absorption coefficients used.

A final observation is that many of the simple models have been based on actual measured data rather than comparison to rigorous models. This fact can make a difference in the direct normal irradiance or the diffuse irradiance but should not affect the total irradiance. The reason for this is that pyrheliometers measure the irradiance in a 5.8-degree field-of-view, which includes some diffuse or circumsolar irradiance. The rigorous codes include only the direct normal irradiance with no circumsolar. This means that the direct normal irradiance with the Bird model will slightly underestimate the irradiance measured by a pyrheliometer. On a normal clear day, one is talking about less than a 1% underestimation. Let us reiterate that the total insolation should agree.



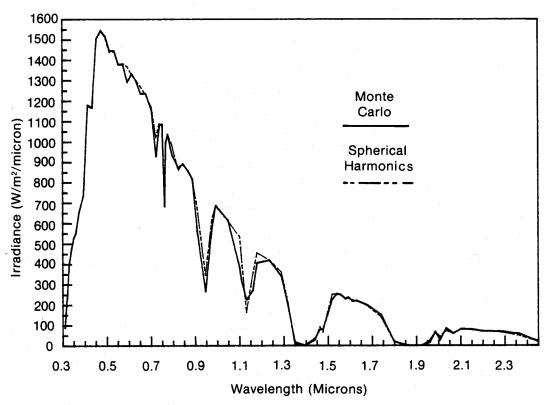


Figure 3-18. Direct Normal Insolation — Dave Model 3 (AM1)

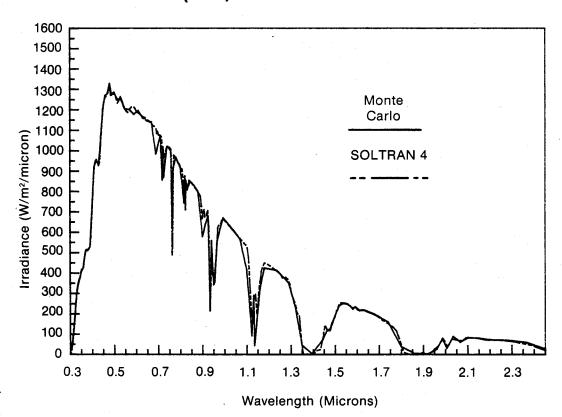


Figure 3-19. Direct Normal Insolation—USS Atmosphere (23-km-Visibility Rural Aerosol, AM1)



#### SECTION 4.0

#### SUMMARY AND CONCLUSIONS

Five simple broadband models for clear sky global horizontal insolation have been compared with the spectrally integrated results from three rigorous spectral codes. As a part of this comparison, a sixth simple broadband model has been formulated. This sixth model, designated the Bird model, uses parts of the formalisms from the other simple models and has been fine-tuned to provide good agreement with the rigorous codes. The Bird model was constructed so that readily available meteorological data could be used in it. It is based entirely on algebraic expressions rather than look-up tables, which greatly simplifies the use of the model.

The comparison of the results from each of the simple models with the results of the rigorous codes indicates the following:

- The Atwater and Ball model is applicable to extremely clear atmospheric conditions with an atmospheric turbidity (base e) near 0.1- at 0.5-µm wavelength. For turbidities near 0.27, this model underestimated the global irradiance by approximately 8% for air mass 1 (AM1). This model is extremely simple but does not have a good method of treating aerosol transmittance.
- The Watt model is relatively complicated and appears to overestimate the global insolation for AMI conditions by approximately 7%. This is a complete model based on meteorological parameters. However, the upper air turbidity required in this model is not readily available.
- The Hoyt model provides excellent agreement with the rigorous codes. However, its use of look-up tables and the requirement to recalculate transmittance and absorptance parameters for modified air mass values causes this model to be relatively difficult to use.
- The Lacis and Hansen model is extremely simple. It tends to overestimate the global irradiance by approximately 8% at AM1, and it has no provisions for calculating direct irradiance.
- The Davies and Hay model could possibly provide good agreement with the rigorous codes. However, it uses a look-up table for the Rayleigh scattering transmittance term and does not have a good method for treating aerosol transmittance. The aerosol transmittance through a vertical path used by Davies and Hay for southern Ontario (K = 0.91) is for an extremely clear atmosphere.
- It is hoped that the rigorous codes and accurate simple models will provide results that will agree within ± 5% with quality experimental data on clear days. Cloudy days are much more difficult to model accurately, and clouds can have the greatest effect on the total irradiance. Models that address cloud influences for irradiance will be examined at a later date.



• It should be recalled that the basis of comparison/evaluation of the simple models is the much more rigorous radiative transfer codes—as opposed to a comparison with actual data. Because of a lack of suit—able, high-quality data, comparisons with actual data are impossible at this time. The greatest deficiency has been the lack of meteorological measurements accompanying good insolation data. However, efforts\* are currently underway at SERI and several universities to provide such data. As this data becomes available, comparisons and improvements will be made. Until then, it appears that both the Hoyt and Bird simplified models yield results in good agreement with the rigorous techniques. However, the Bird model is more flexible and easily used.

<sup>\*</sup>As part of the U.S. Department of Energy's Insolation Resource Assessment Program.



#### SECTION 5.0

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# APPENDIX A

# TABULATED MODEL DATA



Three sets of tabulated data from each simple model are presented here, and each set represents one of the atmospheric models discussed in the main text. The parameters listed at the top of each table are:

 $I_0$  = Solar constant  $(W/m^2)$ 

UW = Precipitable water vapor in vertical path (cm)

UO = Ozone amount in vertical path (cm)

PR = Surface pressure (mb)

TAU5 = Turbidity at  $0.5-\mu m$  wavelength

TAU38 = Turbidity at  $0.38-\mu m$  wavelength

RS = Ground albedo

TEMP = Surface temperature (K)

CONST = Constant K used in BIRD model

BA = Forward to total scattered irradiance ratio

TAUB = Broadband turbidity

The parameters at the top of each column of data are self-explanatory for the transmittance and absorptance terms. The remaining parameters are:

11.

DIRH = Direct horizontal irradiance  $(W/m^2)$ 

DIFSH = Diffuse sky horizontal irradiance  $(W/m^2)$ 

DIFGH = Diffuse ground horizontal irradiance  $(W/m^2)$ 

DTOT = Total or global horizontal irradiance  $(W/m^2)$ 



Table A-1. TABULATED DATA FROM SEVERAL MODELS FOR THE USS ATMOSPHERE

IO = UW = UO = PR = TAUS =	1353 .0000 1 .4200 1 .4200 2400 1013 .0000 2661 3538 .2000 288 .1000 .0733 .8200 .1907 ZENI 0 .00 20 .00 30 .00 48 .19 50 .00 60 .00	00 00 00 00 00	AIRMASS 9995 1.0634 1.1536 1.4972 1.5525 1.9927						
	70.00 75.00 80.00	00	2.8997 3.8076 5.5790						
	85.00	100	10.3163						
		ATI	WATER MO	DEL					
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0 80.0	TA 8264 8163 8024 7513 7434 6832 5736 4806 3382 1249	AW 0855 0872 0873 0966 0977 1053 1179 1281 1441	TMD 9381 9355 9319 9195 9176 9038 8795 8585 8233	100. 40.	2846 6250 1165 4300 1187	64 62 60 51 50 41 30 22	.8348 :	12.2298 11.0037 7.7461 7.3637	DTOT 966.4600 892.6897 803.1906 565.4095 337.4954 374.2064 204.9324 124.6335 54.7216 8.5433
		WA	TT MODEL						
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 75.0 85.0	TA TH20A 8653 ,9250 8590 ,9241 8503 ,9229 8187 ,9192 8138 ,9186 7770 ,9150 70792 ,9096 6544 ,9056 5620 ,9000 3961 ,8904	TH208 9694 9675 9648 9545 9528 9398 9132 8873 8373	. 9749 . 9723	TAIR .9016 .8974 .8917 .8714 .8683 .8451 .8029 .7702 .7142 .6123	923 854 771 548 522 369 210 135 67	888 205 932 831 969 444	DIFSH 176.483 171.361 165.065 147.788 145.694 133.040 118.758 111.100 103.151 94.951	DIFGH 11.272 10.680 9.957 7.996 6.352 4.789 3.963 3.117 2.255	1036.929 946.227 704.717 676.286 509.361 333.991 250.659
		но	YT MODEL						
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0	TAS 8317 8220 8084 7588 7511 6926 5859 4956	AA 0416 0411 0404 0379 0376 0346 0293	AW .1000 .1021 .1049 .1144 .1158 .1258 .1422		ACO2 0075 0076 0078 0084 0085 0091 0101		A03 .0268 .0275 .0285 .0317 .0322 .0357 .0417 .0466	A02 .0075 .0079 .0085 .0107 .0110 .0137 .0190	TR 9170 9125 9063 8840 8807 8560 8138
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0	DIRH 842.7127 776.0145 695.2566 482.1123 457.2772 313.2352 167.1730 99.6955	DIFSH 185.27 183.38 180.81 171.67 170.28 159.86 141.53	187 1 179 1 148 1 128 1 122 1 184	DIFGH 9.8473 8.4825 6.8257 2.4167 1.8978 8.8496 5.6278 4.0258		DT0 047.8 977.8 892.8 666.2 639.4 481.9 314.3	387 849 972 018 573 533		



Table A-1. TABULATED DATA FROM SEVERAL MODELS FOR THE USS ATMOSPHERE (concluded)

		LACIS	MODEL			
ZEN 0.0 20.0 30.0	AW .1089 .1109	A03 . 0234 . 0240 0249	DTOT 1134.0234 1059.6338 969.0372			
48.2 50.0 60.0	.1220 .1233 .1319	.0283 .0288 .0327 .0403	725.7510 696.8444 525.3496 339.5933			
75.0 80.0 85.0	.1559 .1711 .1969	.0474 .0602 .0906	244.5878 150.9857 63.4795			
		DAVIE	S MODEL			
ZEN 0.0 20.0 30.0 48.2 50.0 70.0 75.0 80.0 85.0	DIRH 964.6041 893.1822 806.4406 575.4248 548.2303 388.6384 221.5297 140.8761 68.5779 16.1786	DIFSH 131.3193 129.3610 126.7305 117.7340 116.40771 06.7716 90.5812 77.6703 58.6164 29.9684	DIFGH 18.2102 17.1608 15.8783 12.3977 11.9791 9.4598 6.6085 5.0465 3.3589 1.5249	DTOT 1114.1335 1039.7040 949.0494 705.5564 676.6171 504.8698 318.7193 223.5929 130.5533 47.6719		
		BIRD	MODEL			
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0 85.0	TA 8127 8029 .7895 .7410 .7336 .6778 .5785 .4959 .3704	T03 .9822 .9814 .9803 .9763 .9756 .9709 .9619 .9637 .9391	9874 9 9872 9 9869 9 9860 8 9859 8 9849 8 9834 8 9832 7 9803 7	TR TAS 137 8271 194 8180 1033 8055 116 7603 7541 531 7029 1074 6126 1084 5435 1078 4494 157 4256	AW .1032 .1048 .1068 .1134 .1144 .1249 .1312 .1389 .1499 .1684	TAA .9825 .9815 .9801 .9746 .9728 .9643 .9444 .9124 .8242 .4130
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 75.0 80.0	DIRH 844.2037 777.9391 697.7136 485.8908 461.1879 317.6805 171.6204 103.9955 46.5685 9.0703	DIFSH 168.9023 166.1341 162.3802 149.7404 147.4221 133.6742 110.5682 90.4421 59.8652 12.5752	DIFGH 20.5954 19.5137 18.1816 14.5168 14.0423 11.2868 8.0237 6.0411 3.6917 .7707	DTOT 1033 7014 963 5869 878 2754 650 1480 622 6523 462 6415 290 2122 200 4786 110 1254 22 4162		



Table A-2. TABULATED DATA FROM SEVERAL MODELS FOR THE MLS ATMOSPHERE

1353.0000

UW = UW = UO = PR = TAUS = TAU38 = RS = TEMP = CONST = BA = TAUB =	2.9300 2.9300 3100 1013.0000 .2661 .3538 .8000 .0933 .8200 .1907	·	<b>TDV-05</b>							
	ZEN 0.0 20.0 30.0 48.1 50.0 60.0 75.0 80.0	000 000 000 900 900 000 000 000	IRMASS .9995 1.0634 1.1536 1.4972 1.5525 1.9927 2.8997 3.8076 5.5790	•						
		AT	IATER MOI	DEL.						
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 75.0 85.0	TA 8264 8163 8024 7513 7434 6832 5736 4806 3382 1249	AW 1063 1083 1110 1200 1214 1308 1465 1592 1790 2177	TMD 9381 9355 9319 9195 9176 9038 8795 8585 8233	37.	0702 5702 7362 5418 7946 5108	64 60 51 51 30 22	FSH 4.9349 2.8348 0.0704 1.2334 0.0130 1.7589 0.0717 2.3903 3.2023 3.5038	53 49 44 31 29 20 11	IFGH .9190 .7737 .7471 .4106 .8471 .7135 .2782 .8235 .9680 .4522	DTOT 983 9241 908 2787 816 5538 573 1858 544 6547 377 9832 205 8069 124 5168 54 1614 8 2521
		WA"	IT MODEL							
ZEN 0.0 20.0 30.0 48.2 50.0 75.0 80.0	TA TH2DA 8836 9146 8779 9137 8701 9125 8417 9188 8373 9183 8038 9047 7413 8992 6904 8953 6032 8896	9379 9341 9287 9083 9051 8797 8291 7813 6933	.9763 .9757 .9733 .9729 .9698 .9635 .9586	TAIR 9016 8974 8917 8714 8683 8451 8029 7702 7142 6123	903. 834. 751. 531. 505. 354. 197. 124. 59.	679 725 559 732 681 827	DIF 220 214 206 184 182 166 148 138 129 118	790 383 506 892 271 441 573 992 047	DIFGH 58.727 55.738 52.083 42.168 40.978 33.849 25.750 17.457	3 1104.800 3 1010.315 3 758.619 8 728.982 9 554.970 2 372.332 0 285.586 7 205.740
		но	YT MODEL							
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0	TAS 8317 8220 8084 7588 7511 6926 5859	AA .0416 .0411 .0404 .0379 .0376 .0346 .0293	AW .1272 .1298 .1333 .1451 .1469 .1592 .1796		ACO2 0075 0076 0078 0084 0085 0091 0101		A03 0258 0264 0274 0305 0310 0344 0401		A02 .0075 .0079 .0085 .0107 .0110 .0137 .0190	TR .9170 .9125 .9063 .8840 .8807 .8560 .8138 .7781
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0	DIRH 815 - 7084 750 - 5957 671 - 8142 464 - 2596 440 - 1194 300 - 3540 159 - 2519 94 - 4394	DIFSH 179.34 177.38 174.71 165.31 163.89 153.29 134.82	15 7 10 6 81 6 57 4 30 4 41 3 42 2	DIFGH 2.9219 7.8237 1.6398 5.2159 3.2869 1.9779 0.0839		995. 908. 574. 647. 485.	0T 9718 8004 1721 7908 2989 6256 1600 4540			



Table A-2. TABULATED DATA FROM SEVERAL MODELS FOR THE MLS ATMOSPHERE (concluded)

		LACIS	MODEL				
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0 80.0	AW 1327 1350 1379 1475 1489 1585 1736 1849 2012	A03 0224 0231 0239 0270 0275 0312 0382 0448 0568 0851	DTOT 1138.091: 1063.022: 971.631: 726.431: 697.322: 524.767: 338.225: 243.037: 149.494: 62.426:	3 1 0 1 9 0 3			
		DAVIE	S MODEL				
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 80.0 85.0	DIRH 936.3331 866.5097 781.7525 556.3066 529.8009 374.4382 212.2479 134.2798 64.7464 14.9323	DIFSH 129.0411 127.0680 124.4193 115.3767 114.0456 104.3899 88.2198 75.3692 56.4768 28.3066	DIFGH 74.525; 70.234' 64.993; 50.779; 49.070; 38.794' 27.173; 20.805 13.906' 6.344	3 1139 9 1063 3 971 0 722 8 692 9 517 9 327 3 230 7 135	FOT 8995		
		BIRD	MODEL				
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0 80.0	TA 8127 8029 7895 7410 7336 6778 5785 4959 3704	T03 .9834 .9826 .9816 .9778 .9772 .9727 .9643 .9566 .9430	.9860 .9859 .9849	TR .9137 .9094 .9033 .8816 .8783 .8531 .8074 .7684 .7078 .6157	TAS 6271 8180 8055 7603 7541 7029 6126 5435 4494	AW .1219 .1235 .1257 .1330 .1340 .1411 .1520 .1601 .1717 .1907	TAA 9825 9815 9801 9746 9728 9643 9444 9124 8242
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 75.0 80.0 85.0	DIRH 827.6234 762.5502 683.7850 475.9279 451.6994 311.0063 167.9332 101.7433 45.5637 8.8855	DIFSH 165.5850 162.8477 159.1385 146.6700 144.3891 130.8658 108.1927 88.4834 58.5735 12.3191	DIFGH 86.007 81.568 76.100 61.060 57.104 47.783 34.333 26.071 16.126 3.381	1 1079 9 1006 7 919 3 683 8 655 4 489 8 310 0 216 3 120	707 .2176 .9668 .0243 .6583 .1932 .6554 .4597 .2978 .2635 .5857		



10 = 1353.0000

Table A-3. TABULATED DATA FROM SEVERAL MODELS FOR THE DAVE MODEL 3 ATMOSPHERE

1U = UU = TOUS = TAUS = TAU38= TEMP = TOUST= TAUB =	29300 3100 1013.0000 1013.0000 10979 .0979 .0270 294.0000 .0933 .8600 .0620 ZEN 0.0 20.0 30.0 48.1 50.0 60.0 75.0 85.0	000 000 000 900 000 000 000 000	AIRMASS .9995 1.0634 1.1536 1.4972 1.5525 1.9927 2.8997 2.8997 3.5790 10.3163						
		ATI	WATER MO	DEL.					•
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0 80.0	TA 9399 9362 9310 9113 9081 8836 8348 7881 7031	AW .1063 .1083 .1110 .1200 .1214 .1308 .1465 .1592 .1790 .2177	TMD .9381 .9355 .9319 .9195 .9176 .9038 .8795 .8585 .8233 .7472		9331 4993 7991 9837 8031 0431	73 72 69 62 61 54 43 36	FSH .8571 .0609 .6973 .1416 .0969 .0058 .7662 .7179 .4474 .2687	DIFGH 1.4512 1 1.3507 1.2285 9015 8628 6339 3884 2648 1460 0436	DTOT 1059 2413 985 9109 896 7250 658 0268 629 7628 462 6827 283 5044 193 2703 106 5762 31 8077
		WA	TT MODEL						
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 85.0	TA TH20A 9762 9144 9743 9137 9716 9125 9609 9088 9452 9083 9452 9084 9157 8992 8921 8955 8457 8896 7602 8800	9379 9341 9287 9083 9083 9051 98797 88797 88291 87813 6933	T03 .9768 .9763 .9757 .9733 .9729 .9698 .9635 .9586 .9489 .9309	TAIR 9016 8974 8917 8714 8683 8451 8029 7702 7142 6123	997. 926. 839. 606. 579. 417. 244. 161.	307 372 839 335 053	DIFSH 158.878 154.267 148.600 133.046 131.160 119.769 106.911 100.017 92.861 85.480	DIFGH 1 029 97/ 901 721 70 57/ 421 35:	4 1081.548 7 988.879 5 740.610 3 711.198 2 537.395 8 351.717 1 261.680 3 176.179
		но	YT MODEL						
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0	TAS .9321 .9279 .9220 .9000 .8965 .8692 .8155	AA 0466 0464 0461 0450 0448 0435 0438	AW .1272 .1298 .1333 .1451 .1469 .1592 .1796		AC02 0075 0076 0078 0084 0085 0091 0101	-	A03 .0258 .0264 .0274 .0305 .0310 .0344 .0401	A02 .0075 .0079 .0085 .0107 .0110 .0137 .0190	TR .9170 .9125 .9063 .8840 .8807 .8560 .8138
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0	DIRH 908.3464 841.6283 760.6718 545.5922 520.3379 372.5084 218.1138 142.9609	DIFSH 98.20 97.21 95.87 91.16 90.45 85.16 76.09 68.97	25 54 85 94 80 41	DIFGH 1.2750 1.1857 1.0773 .7888 .7548 .5554 .3449 .2407		DTO 007.8 940.0 357.6 637.5 611.5 458.2 294.5	240 295 276 504 508 279		



Table A-3. TABULATED DATA FROM SEVERAL MODELS FOR THE DAVE MODEL 3 ATMOSPHERE (concluded)

		LACIS	MODEL			
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 75.0 80.0	AW 1327 1350 1379 1475 1489 1585 1736 1849 2012	A03 .0224 .0231 .0239 .0270 .0275 .0312 .0382 .0448 .0568	DTOT 1073.1443 1020.9793 933.1311 697.4798 609.5094 503.7295 324.5709 233.1836 143.4038 59.8739			
		DAVIE	S MODEL			
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 75.0 85.0	DIRH 936.3331 866.5097 781.7525 556.3066 529.8009 374.4382 212.2479 134.2798 64.7464 14.9323	DIFSH 129.0411 127.0680 124.4193 115.3767 114.0456 104.3899 88.2198 75.3692 56.4768 28.3066	DIFGH 1.7442 1.6427 1.5186 1.1823 1.1419 .8989 .6243 .4742 .3127	DTOT 1067.1183 995.2204 907.6905 672.8657 644.9885 479.7270 301.0919 210.1232 121.5359 43.3777		
		BIRD	MODEL			
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 75.0 80.0 85.0	TA 9219 91175 9115 8891 8856 8585 8068 7595 6774 5057	T03 9834 9826 9816 9778 9772 9727 9643 9566 9430	.9874 .9 .9872 .9 .9869 .9 .9860 .8 .9859 .8 .9849 .8 .9834 .8 .9832 .7 .9803 .7	TR TAS 137 9286 094 9247 033 9192 816 8989 783 8961 531 8722 074 8279 684 7926 078 7444 157 7804	AW 1219 1235 1257 1330 1340 1411 1520 1601 1717	TAA 9927 9923 9916 9891 9883 9843 9745 9582 9099
ZEN 0.0 20.0 30.0 48.2 50.0 60.0 70.0 75.0 80.0 85.0	DIRH 938.8409 871.3803 789.4391 571.0358 545.2946 393.9366 234.2099 155.8315 83.3321 25.5650	DIFSH 74.5702 93.1824 91.3142 85.0776 83.9642 77.2467 65.8789 55.7656 39.5922 11.1038	DIFGH 1.624B 1.5273 1.4082 1.0865 1.0470 .8155 .5567 .4135 .2569	DTOT 1035.0359 966.0901 882.1616 657.1999 630.3058 471.9989 300.6456 212.0107 123.1812 36.7417		

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global insolati Watt, Hoyt, and been formulated results from th vide results th in the zenith p and with the re	arison was made between on. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are	were those of Atward A sixth simple models and we transfer codes. with the three rigord Hoyt models agrees codes. However,	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to
A detailed comp global insolati Watt, Hoyt, and been formulated results from th vide results the in the zenith p and with the re	arison was made between on. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous	were those of Atward A sixth simple models and we transfer codes.  with the three rigord Hoyt models agrees codes. However,	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to
A detailed comp global insolati Watt, Hoyt, and been formulated results from th vide results the in the zenith p and with the re	arison was made between on. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous	were those of Atward A sixth simple models and we transfer codes.  with the three rigord Hoyt models agrees codes. However,	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to
A detailed comp global insolati Watt, Hoyt, and been formulated results from th vide results the in the zenith p and with the re	arison was made between on. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous	were those of Atward A sixth simple models and we transfer codes.  with the three rigord Hoyt models agrees codes. However,	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to
A detailed comp global insolati Watt, Hoyt, and been formulated results from th vide results the in the zenith p and with the re	arison was made between on. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous	were those of Atward A sixth simple models and we transfer codes.  with the three rigord Hoyt models agrees codes. However,	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to
A detailed comp global insolati Watt, Hoyt, and been formulated results from th vide results the in the zenith p and with the re	arison was made between on. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous	were those of Atward A sixth simple models and we transfer codes.  with the three rigord Hoyt models agrees codes. However,	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to
A detailed comp global insolati Watt, Hoyt, and been formulated results from th vide results the in the zenith p and with the re	arison was made between on. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous	were those of Atward A sixth simple models and we transfer codes.  with the three rigord Hoyt models agrees codes. However,	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to
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A detailed comp global insolati Watt, Hoyt, and been formulated results from the vide results the in the zenith pand with the resimplement and half the complement and half th	arison was made between. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous as broader application.	were those of Atward A sixth simple model and the service of the sixth simple models and the service of the ser	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the. All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to del.  s; Insolation:Tl; Mathematic.
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A detailed comp global insolati Watt, Hoyt, and been formulated results from the vide results the inthe zenith pand with the resimplement and half and been formulated results from the zenith pand with the resimplement and half and with the resimplement and half and be solved by the senior of the	arison was made between. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous as broader application. The sults of the rigorous as broader application. The sults of the rigorous are broader application. The sults of the rigorous are broader application.	were those of Atward A sixth simple model and the service of the sixth simple models and the service of the ser	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the. All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to del.  s; Insolation:Tl; Mathematic.
A detailed comp global insolati Watt, Hoyt, and been formulated results from the vide results the inthe zenith pand with the resimplement and half the senith pand with the s	arison was made between. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% cosition. The Bird are sults of the rigorous as broader application. The sults of the rigorous as broader application. The sults of the rigorous are broader application. The sults of the rigorous are broader application.	were those of Atward A sixth simple model and the service of the sixth simple models and the service of the ser	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the. All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to del.  s; Insolation:Tl; Mathematic.
A detailed comp global insolati Watt, Hoyt, and been formulated results from the vide results the inthe zenith pand with the resimplement and hard with the resimplement and hard bearing and bearing	arison was made between. Compared models Lacis and Hansen. A by using parts of the ree rigorous radiative at agree within <10% rosition. The Bird are sults of the rigorous as broader application.  Imparative Analysis; asuring Methods; Methoded Terms	were those of Atward A sixth simple model and the service of the sixth simple models and the service of the ser	ter and Ball, Davies and Hay, l, called the Bird model, has d by comparison with the. All of the simple models proprous codes when the sun is e vithin 3% with each other the Bird model is easier to del.  s; Insolation:Tl; Mathematic.
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