



NREL Pyrheliometer Comparisons: September 25–October 6, 2017 (NPC-2017)

Ibrahim Reda, Mike Dooraghi, Afshin Andreas,
and Aron Habte
National Renewable Energy Laboratory

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Figure 1. NPC-2017 participants

List of Acronyms

AHF	Automatic Hickey-Frieden
BMS	Baseline Measurement System
BORCAL	Broadband Outdoor Radiometer Calibration
DOE	U.S. Department of Energy
IPC	International Pyrheliometer Comparison
IPC-XII	Twelfth International Pyrheliometer Comparisons
ISO	International Organization for Standardization
MST	Mountain Standard Time
NPC	National Renewable Energy Laboratory Pyrheliometer Comparisons
NREL	National Renewable Energy Laboratory
PMOD/WRC	Physikalisch-Meteorologisches Observatorium Davos—World Radiation Center
SDp	pooled standard deviation
SI	International System of Units
SRRL	Solar Radiation Research Laboratory
TSG	Transfer Standard Group
WMO	World Meteorological Organization
WRR	World Radiometric Reference
WRR-TF	World Radiometric Reference transfer factor
WSG	World Standard Group
S/N	serial number of radiometer
%uA	Percentage Type-A standard uncertainty
NRdg	number of readings
uC	combined standard uncertainty
Eff DF	effective degrees of freedom

Executive Summary

Accurate measurements of direct normal (beam) solar irradiance from pyrheliometers¹ are important for developing and deploying solar energy conversion systems, improving our understanding of the Earth's energy budget for climate change studies, and for other science and technology applications involving solar flux. Providing these measurements places many demands on the quality system used by the operator of commercially available radiometers. Maintaining accurate radiometer calibrations that are traceable to an international standard is the first step in producing research-quality solar irradiance measurements.

In 1977, the World Meteorological Organization (WMO) established the World Radiometric Reference (WRR) as the international standard for the measurement of direct normal solar irradiance (Fröhlich 1991). The WRR is an internationally recognized, detector-based measurement standard determined by the collective performance of six electrically self-calibrated absolute cavity radiometers comprising the World Standard Group (WSG). Various countries, including the United States,² have contributed these specialized radiometers to the Physikalisch-Meteorologisches Observatorium Davos—World Radiation Center (PMOD/WRC) to establish the WSG.

As with all measurement systems, Absolute Cavity Radiometers (ACR) are subject to performance changes over time. Therefore, every five years the PMOD/WRC in Davos, Switzerland, hosts an International Pyrheliometer Comparison (IPC) for transferring the WRR to participating radiometers. NREL has represented the U.S. Department of Energy (DOE) in each IPC since 1980. As a result, NREL has developed and maintained a select group of absolute cavity radiometers with direct calibration traceability to the WRR, and uses these reference instruments to calibrate pyrheliometers and pyranometers using the International Organization for Standardization (ISO) 17025 accredited Broadband Outdoor Radiometer Calibration (BORCAL) process (Reda et al. 2008).

National Renewable Energy Laboratory (NREL) pyrheliometer comparisons (NPCs) are held annually at the Solar Radiation Research Laboratory (SRRL) in Golden, Colorado. Open to all ACR owners and operators, each NPC provides an opportunity to determine the unique WRR transfer factor (WRR-TF) for each participating pyrheliometer. By adjusting all subsequent pyrheliometer measurements by the appropriate WRR-TF, the solar irradiance data are traceable to the WRR.

NPC-2017 was September 25 through October 6, 2017. Participants operated 45 ACRs to simultaneously measure clear-sky direct normal solar irradiance during this period. The Transfer Standard Group (TSG) of reference radiometers for NPC-2017 consisted of four NREL radiometers with direct traceability to the WRR, having participated in the Twelfth International Pyrheliometer Comparisons (IPC-XII) in the fall of 2015. As a result of NPC-2017, each participating absolute cavity radiometer was assigned a new WRR-TF, computed as the reference irradiance determined by the TSG divided by the observed irradiance from the participating radiometer. The performance of the TSG during NPC-2017 was consistent with previous comparisons, including IPC-XII. The measurement performance of the TSG allowed

¹ Pyrheliometers are a type of radiometer used to measure solar irradiance (i.e., radiant flux in Watts per square meter) on a surface normal to the apparent solar disk within a 5.0° or 5.7° field of view, depending on the optical design of the instrument. A solar tracker is used to maintain proper alignment of the pyrheliometer with the sun during daylight periods.

² The WSG includes radiometers on permanent loan from the Eppley Laboratory, Inc., and NREL.

the transfer of the WRR to each participating radiometer with an estimated uncertainty of $\pm 0.36\%$ with respect to the International System of Units.

The comparison protocol is based on data collection periods called *runs*. Each measurement run consists of an electrical self-calibration requiring five minutes for the Automatic Hickey-Frieden (AHF) cavities, a series of 49 solar irradiance measurements at 30-second intervals, and a post-calibration. More than 2,000 reference irradiance measurements were collected by the TSG during NPC-2017. Clear-sky daily maximum direct normal irradiance level was 996 Wm^{-2} .

Ancillary environmental conditions (e.g., broadband turbidity, ambient temperature, relative humidity, wind speed, precipitable water vapor, and spectral data) collected at SRRL during the comparison are presented in Appendix B to document the environmental test conditions.

NPCs are planned annually at the SRRL to ensure worldwide homogeneity of solar radiation measurements traceable to the WRR.

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1 Introduction

Accurate measurements of broadband solar irradiance require radiometers with proper design and performance characteristics, correct installation, and documented operation and maintenance procedures, including regular calibration. Calibrations of any measuring device must be traceable to a recognized reference standard. The World Radiometric Reference (WRR) is the internationally recognized measurement standard for direct normal irradiance measurements of broadband solar radiation (Fröhlich 1991).

The WRR was established by the World Meteorological Organization (WMO) in 1977 and has been maintained by the Physikalisch-Meteorologisches Observatorium Davos—World Radiation Center (PMOD/WRC) in Switzerland (www.pmodwrc.ch). This reference is maintained for broadband solar irradiance with an absolute uncertainty of better than $\pm 0.3\%$ with respect to the International System of Units (SI) (Romero et al. 1996). This standard is widely used to calibrate pyrheliometers and pyranometers with a wavelength response range that is compatible with the solar spectrum wavelengths of 280–3,000 nm. Every five years, the WRR is transferred to WMO regional centers and other participants at International Pyrheliometer Comparisons (IPC) held at the PMOD/WRC. The Twelfth IPC (IPC-XII) was completed in 2015 (Finsterle 2017). At each IPC, instantaneous measurements from the World Standard Group (WSG) are compared at 90-second intervals with the data from participating radiometers recorded under clear-sky conditions. A new WRR transfer factor (WRR-TF) is calculated for each participating radiometer based on the mean WRR of the WSG radiometers for each IPC. Multiplying the irradiance reading of each radiometer by its assigned WRR-TF will result in measurements that are traceable to SI units through WRR and therefore consistent with the international reference of solar radiation measurement.

In compliance with Organization for Standardization (ISO) 17025 accreditation requirements for demonstrating interlaboratory proficiency, the National Renewable Energy Laboratory (NREL) hosts annual pyrheliometer comparisons at the Solar Radiation Research Laboratory (SRRL) in Golden, Colorado, for non-IPC years. The NREL Pyrheliometer Comparisons in 2017 (NPC-2017) was September 25 through October 6, 2017, at the SRRL. Participants operated 40 absolute cavity radiometers during the comparisons. See Appendix A for the list of participants and affiliations.

The results presented in this report are based on clear-sky, direct normal solar irradiance data collected during the NPC. See Appendix B for environmental conditions.

2 Reference Instruments

NREL developed the transfer standard group (TSG) of four absolute cavity radiometers to serve as the transfer reference for each NPC. The radiometers comprising the TSG were included in the most recent IPC and maintain the WRR for NREL (see Table 1). Using the method described by Reda (1996), the mean of the TSG measurements was maintained for establishing the reference irradiance data for NPC-2017 data reduction. Table 1 provides a list of the TSG absolute cavity radiometers with their WRR-TFs and pooled standard deviation (SD_p) as determined from the latest IPC in 2015 (Finsterle 2017).

Table 1. IPC-XII Results Summary for the NPC-2017 TSG

Serial Number	WRR Factor (IPC-XI)	Standard Deviation (%)	Number of Readings
AHF 28968	0.99763	0.063	519
AHF 29220	0.99749	0.0621	523
AHF 30713	0.99723	0.0639	525
ATMI 68018	0.99660	0.0669	522
Mean WRR for the TSG	0.99724	SD_p for the TSG: 0.06%	

The pooled standard deviation, SD_p , for the TSG was computed from the following equation:

$$SD_p = \sqrt{\frac{\sum_{i=1}^m n_i S_i^2}{\sum_{i=1}^m n_i}} \quad 1$$

where,

$i = i^{\text{th}}$ cavity

m = number of reference cavities

S_i = standard deviation of the i^{th} cavity, from IPC-XII

n_i = number of readings of the i^{th} cavity, from IPC-XII

3 Measurement Protocol

The decision to deploy instruments for a comparison was made daily. Data were collected only during clear-sky conditions, which were determined visually and from the stability of pyrheliometer readings. Simultaneous direct normal solar irradiance measurements were taken by most cavity radiometers in groups of 49 observations at 30-second intervals (PMO6 used 80- or 90-second open/closed-shutter cycles). Each group of observations is called a *run*. An electrical self-calibration of each Automatic Hickey-Frieden (AHF) absolute cavity was performed prior to each run. Previous WRR-TFs determined from results of IPCs or NPCs were *not* applied to the observations. The original manufacturer's calibration factor was used according to the standard operating procedure provided by the manufacturer for each radiometer. A timekeeper announced the beginning of each calibration period and gave a five-minute countdown prior to the start of each run to facilitate the AHF cavity self-calibrations and the simultaneous start for each participant.

By consensus, at least 300 observations from each radiometer were required to determine the WRR-TF for an NPC. A statistically significant dataset was required to derive the WRR-TF for each pyrheliometer. Data from each pyrheliometer/operator system were emailed at the end of the day.

4 Transferring the World Radiometric Reference

The primary purpose of an NREL pyrheliometer comparison is to transfer the current WRR from the NPC-TSG to each participating absolute cavity pyrheliometer. This requires that the participating pyrheliometers and the TSG collect simultaneous measurements of clear-sky direct normal (beam) solar irradiance. Because the NPC data analysis is intended for absolute cavity pyrheliometers only, users of pyrheliometers other than absolute cavity pyrheliometers might interpret their NPC results differently.

4.1 Calibration Requirements

Using WMO guidelines (Romero 1995), the following conditions were required before data collection was accomplished during NPC-2017:

- The radiation source was the sun, with irradiance levels $> 700 \text{ Wm}^{-2}$.
- A Digital Multimeter with uncertainty $> 0.05\%$ reading or better was used to measure the thermopile signals from each radiometer.
- Solar trackers were aligned within $\pm 0.25^\circ$ slope angle.
- Wind speed was low ($< 5 \text{ m/s}$) from the direction of the solar azimuth $\pm 30^\circ$.
- Cloud cover was $< 1/8$ of the sky dome, with an angular distance $> 15^\circ$ from the sun.

4.2 Determining the Reference Irradiance

Four absolute cavity radiometers maintained by NREL and were part of IPC-XII were used as the TSG to transfer the WRR in the comparison. The WRR-TF for each TSG is presented in Table 1 above. The reference irradiance at each reading was calculated using the following steps, as described by Reda (1996):

1. Each irradiance reading of the TSG is divided by the irradiance measured by AHF28968, for its participation in many IPCs.
2. By maintaining the mean of WRR for the TSG, a new WRR-TF for NPC-2017 is recalculated for each of the TSG cavities (see Figure 2).
3. The reference irradiance for each 30-second observation in a run is computed as the mean of the simultaneous reference irradiances measured by the TSG. The reference irradiance reading for each cavity in the TSG is the irradiance reading of the cavity multiplied by its new WRR-TF calculated in Step 2.

4.3 Data Analysis Criteria

AHF28968 was used to check irradiance stability at the time of each comparison reading during a run. Stable irradiance readings are defined to within 1.0 Wm^{-2} during an interval of two seconds centered on each reading time (i.e. one second before and one second after the recorded reading). Unstable irradiance readings are marked in the data record and automatically rejected from the data analysis. Historically, this has affected fewer than 10% of the data collected during an NPC.

Additionally, all calculated ratios of the test instrument irradiance divided by AHF28968 irradiance that deviated from their mean by 0.3% were rejected (Reda 1996). Typically, data rejected from the analysis in this manner were the result of failed tracker alignment, problems

with the pre-calibration, or a similar cause for bias greater than expected from a properly functioning absolute cavity radiometer.

4.4 Measurements

NPC-2017 was completed for most participants on October 6, after more than 2,000 data points were collected by the reference cavities during the requisite clear-sky conditions. The actual number of readings for each participating radiometer compared with the reference irradiance varied according to the data analysis selection criteria described above. Additionally, some instruments experienced minor data loss because a variety of problems occurred with the measurement systems and operations.

4.5 Results

The historical results for the TSG are presented in Figure 2. To evaluate the performance of these instruments, the standard deviations of each radiometer were monitored during the comparisons. The results suggest successful performance of the TSG during this NPC:

- For the TSG, the NPC-2017 WRR-TF did not change by more than a fraction of the standard deviation derived during IPC-XII in 2015 (see Figure 2.).
- For NPC-2017 Proficiency Test, the results of the participating cavities in IPC-XII and NPC-2017 were evaluated using the following equation:

$$E_n = \frac{WRR_{IPC} - WRR_{NPC}}{\sqrt{U95_{IPC}^2 + U95_{NPC}^2}} \quad 2$$

where E_n must lie in the interval -1 to +1.

From Table 2, E_n for all cavities was well within the interval -1 to +1 (i.e., the WRR from NPC-2017 is consistent with the WRR from IPC-XII).

WRR-RF for NREL Reference Cavities

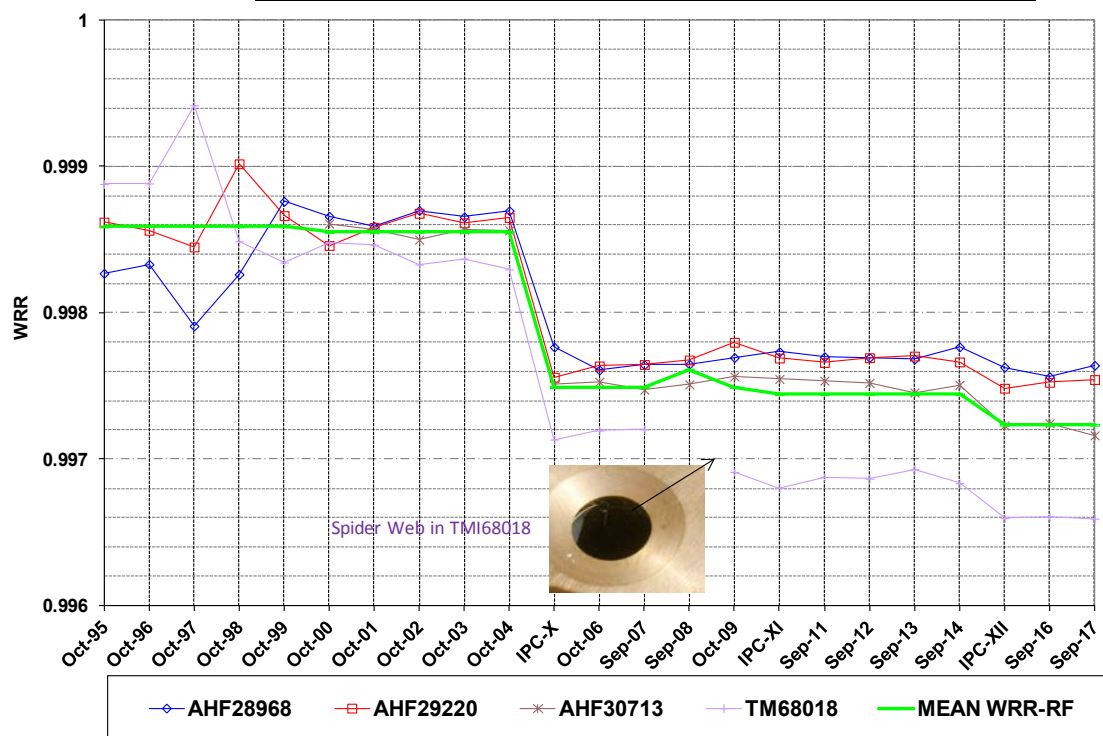


Figure 2. History of WRR reduction factors for NREL reference cavities

Table 2. Summary Results for Proficiency Test During NPC-2017

S/N	WRR (IPC-XII)	%U95 (IPC-XII)	WRR (NPC-2017)	%U95 (NPC-2017)	Red if $\text{abs}(E_n) > 1$
AHF 0000	1.000307	0.36	1.00233	0.39	-0.0038
AHF 14915	0.999542	0.35	0.99916	0.40	0.0007
AHF 14917	0.9979	0.35	0.99775	0.38	0.0003
AHF 17142	0.997946	0.39	0.99828	0.57	-0.0005
AHF 23734	0.998187	0.35	0.99835	0.37	-0.0003
AHF 28553	0.997739	0.35	0.99761	0.38	0.0002
AHF 28556	0.995408	0.36	0.99470	0.38	0.0014
AHF 28560	0.999283	0.39	0.99351	0.45	0.0098
AHF 31041	0.996394	0.36	0.99717	0.47	-0.0013
AHF 31105	0.998657	0.36	0.99880	0.38	-0.0003
AHF 31114AWX	1.001209	0.33	1.00067	0.39	0.0010
AHF 32448AWX	0.999986	0.35	1.00004	0.39	-0.0001
AHF 32455	1.00138	0.33	1.00139	0.38	0.0000
AHF 37816	0.999458	0.33	0.99944	0.39	0.0000
PMO6 0816	0.999947	0.45	1.00010	0.44	-0.0003
PMO6 81109	0.998317	0.32	0.99837	0.39	-0.0001
PMO6 911204	0.999947	0.41	0.99987	0.41	0.0001
Pmo6cc 0103	0.997916	0.32	0.99789	0.40	0.0001
PMO6cc 0401	1.020799	0.31	1.02094	0.38	-0.0003
PMO6-cc 0803	1.000335	0.32	1.00018	0.39	0.0003
TMI 68835	1.000714	0.32	1.00015	0.42	0.0011

Table 3. Results for Radiometers Participating in NPC-2017

S/N	WRR- Reduction Factor (Testcav)	%uA	NRdg	uC	Eff DF	%U95
AHF 0000	1.00233	0.08	1873	0.20	8.73E+04	0.39
AHF 14915	0.99916	0.09	1246	0.21	3.36E+04	0.40
AHF 14917	0.99775	0.05	1283	0.19	2.09E+05	0.38
AHF 17142	0.99828	0.22	1061	0.29	3.03E+03	0.57
AHF 23734	0.99835	0.04	1980	0.19	1.11E+06	0.37
AHF 28553	0.99761	0.05	1291	0.19	2.02E+05	0.38
AHF 28556	0.99470	0.06	1203	0.19	1.13E+05	0.38
AHF 28560	0.99351	0.13	562	0.23	4.75E+03	0.45
AHF 29219-Window	1.06152	0.09	2003	0.20	5.80E+04	0.40
AHF 29222-Window	1.05842	0.06	1668	0.19	2.36E+05	0.38
AHF 30495-Window	1.05512	0.07	1673	0.20	1.06E+05	0.39
AHF 31041	0.99717	0.15	1807	0.24	1.11E+04	0.47
AHF 31104-Window	1.03837	0.06	2004	0.19	2.02E+05	0.38
AHF 31105	0.99880	0.06	1078	0.19	1.12E+05	0.38
AHF 31108	0.99686	0.07	1429	0.20	7.32E+04	0.39
AHF 31111	0.99713	0.06	1295	0.19	1.57E+05	0.38
AHF 31113AWX-Window	1.04975	0.09	1218	0.20	3.81E+04	0.40
AHF 31114AWX	1.00067	0.08	1217	0.20	5.88E+04	0.39
AHF 31116AWX-Window	1.05753	0.09	1230	0.21	3.39E+04	0.40
AHF 32448AWX	1.00004	0.08	1230	0.20	5.39E+04	0.39
AHF 32452AWX-Window	1.03104	0.12	1966	0.22	2.09E+04	0.43
AHF 32455	1.00139	0.06	1852	0.19	2.16E+05	0.38
AHF 33392	0.99897	0.07	1493	0.20	7.85E+04	0.39
AHF 34926AWX	0.99935	0.09	621	0.21	1.65E+04	0.40
AHF 37816	0.99944	0.07	534	0.20	3.86E+04	0.39
PMO6 0816	1.00010	0.13	568	0.22	5.52E+03	0.44
PMO6 1601	1.00308	0.07	662	0.20	3.96E+04	0.39
PMO6 81109	0.99837	0.08	504	0.20	1.99E+04	0.39
PMO6 911204	0.99987	0.10	520	0.21	1.09E+04	0.41
Pmo6cc 0103	0.99789	0.08	521	0.20	1.82E+04	0.40
PMO6cc 0401	1.02094	0.06	451	0.19	5.35E+04	0.38
PMO6-cc 0803	1.00018	0.07	454	0.20	2.99E+04	0.39
TMI 67603	1.00007	0.07	1395	0.20	8.93E+04	0.39
TMI 67811	0.99817	0.07	1362	0.20	8.11E+04	0.39
TMI 68022	0.99933	0.13	1364	0.22	1.31E+04	0.44
TMI 68835	1.00015	0.11	1608	0.21	2.36E+04	0.42

The uncertainty of the WRR-TF associated with each participating radiometer with respect to SI was calculated using the following formula:

$$U_{95} = \pm 1.96 * \sqrt{u_A^2 + u_B^2} \quad 3$$

where,

U_{95} = Uncertainty of the WRR-TF (in percent) determined at NPC-2017 with 95% confidence level

1.96 = Coverage factor

u_A = Type A standard uncertainty = standard deviation of each participating radiometer (in %) determined at NPC-2017

u_B = Type B standard uncertainty

$$u_B = \pm \sqrt{\left(\frac{0.3}{\sqrt{3}}\right)^2 + 0.06^2}$$

where,

0.3 = Estimated expanded uncertainty of the WRR scale with respect to SI, in %

$\sqrt{3}$ = Coverage factor for rectangular distribution

0.06 = Pooled standard deviation of the four reference radiometers (TSG) that participated in IPC-XII (September/October 2015), in %.

The statistical analyses of WRR-TF for the participating pyrheliometers are presented in the following figures. These graphical summaries indicate the mean, standard deviation, and histograms of the WRR-TF determined during NPC-2017.

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

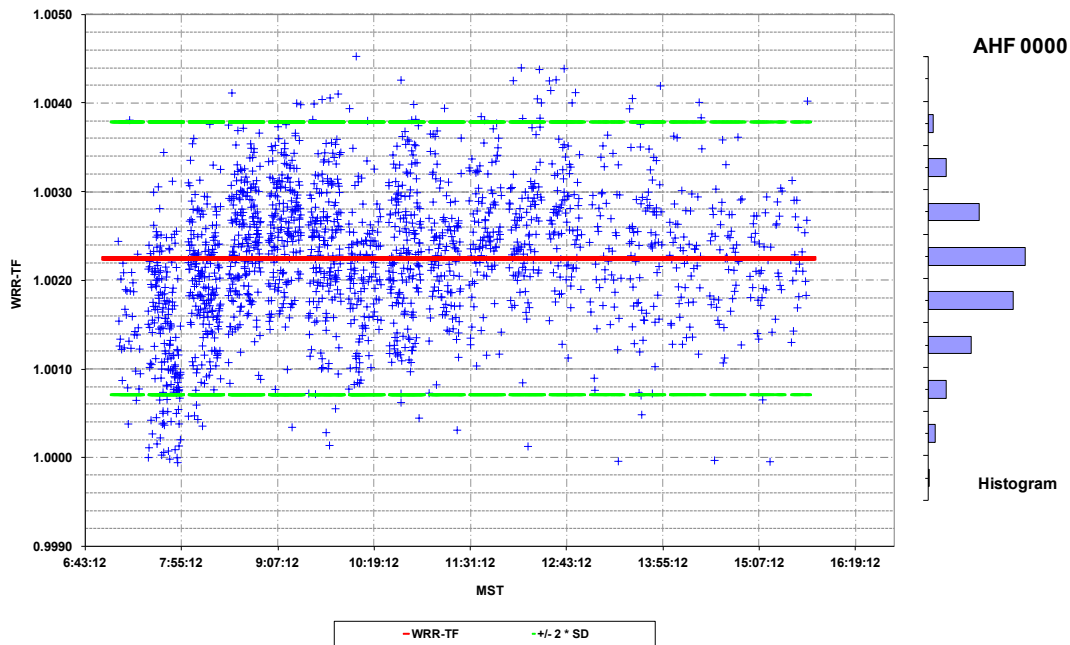


Figure 3. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 0000

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

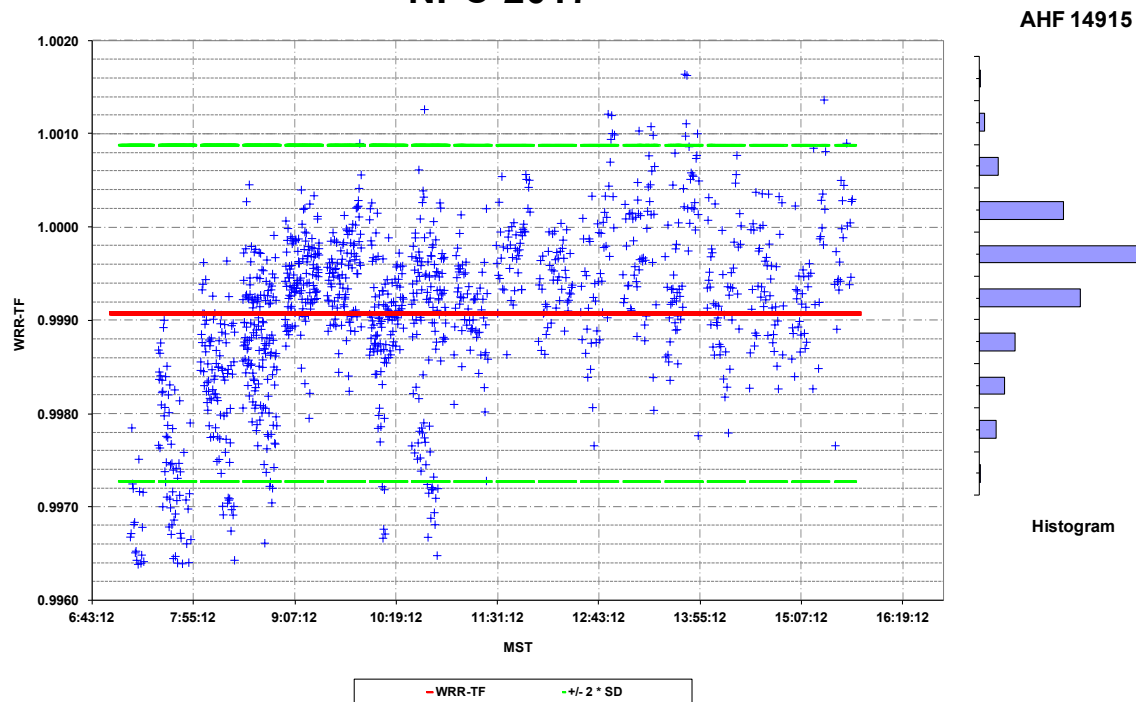


Figure 4. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 14915

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

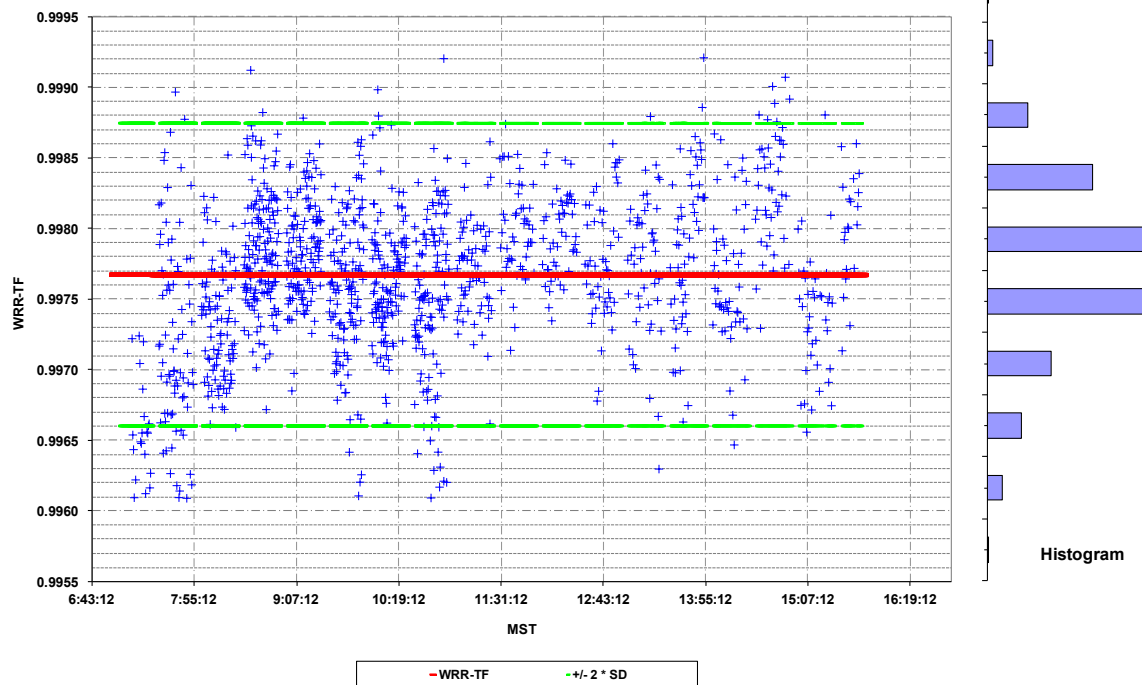


Figure 5. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 14917

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

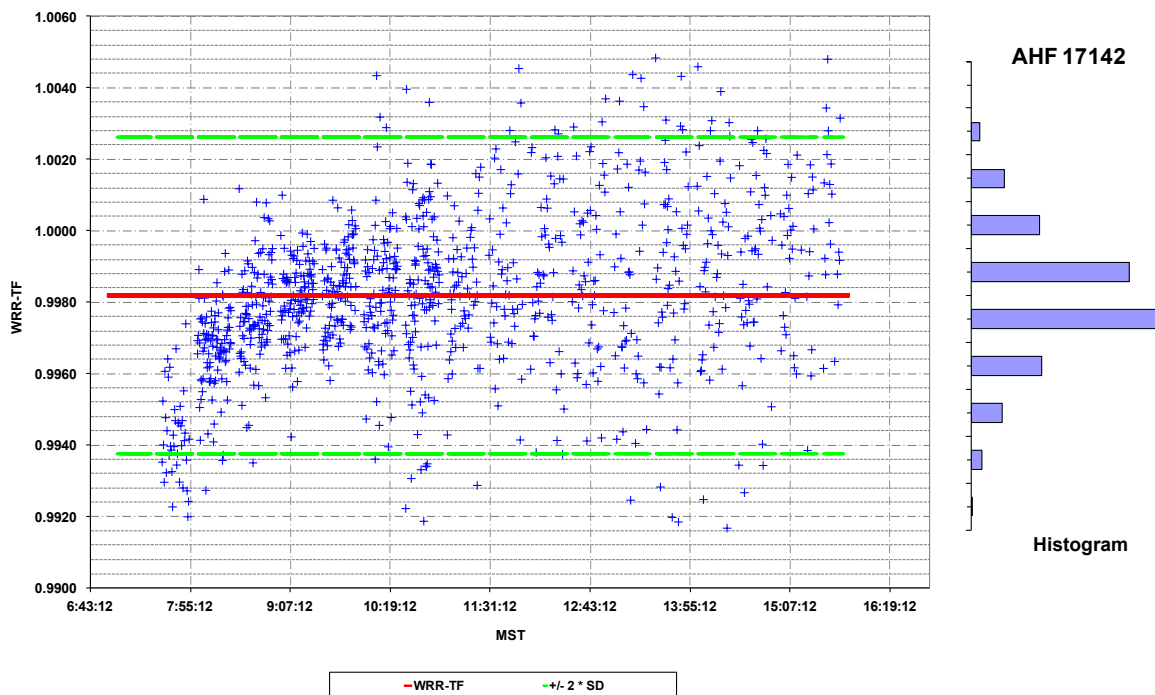


Figure 6. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF17142

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

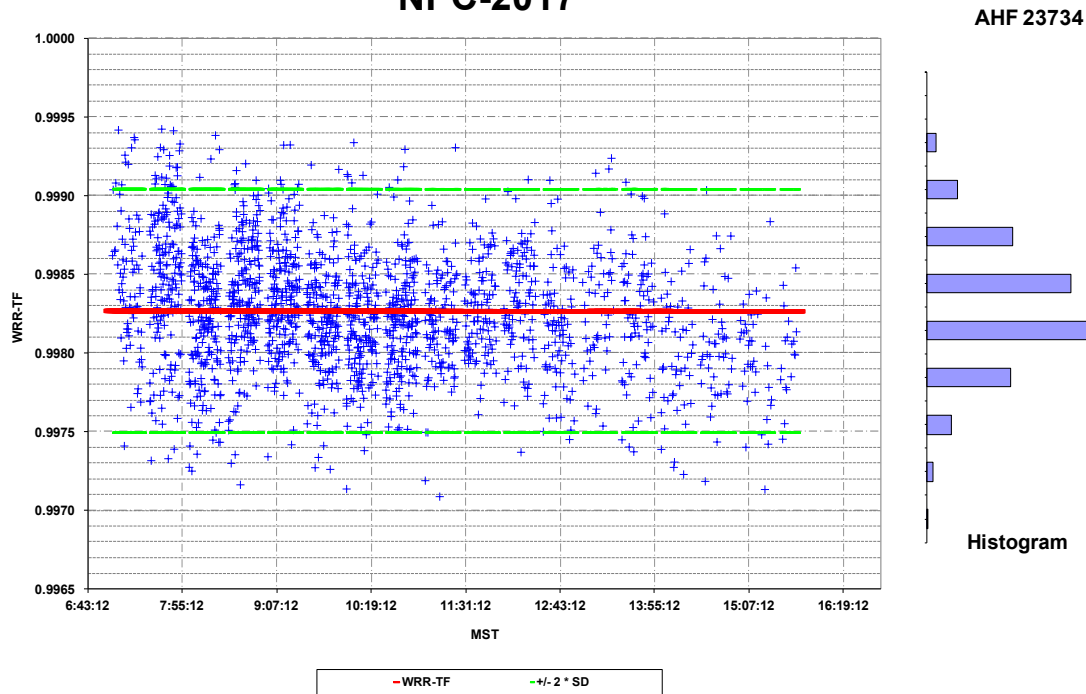


Figure 7. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 23734

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

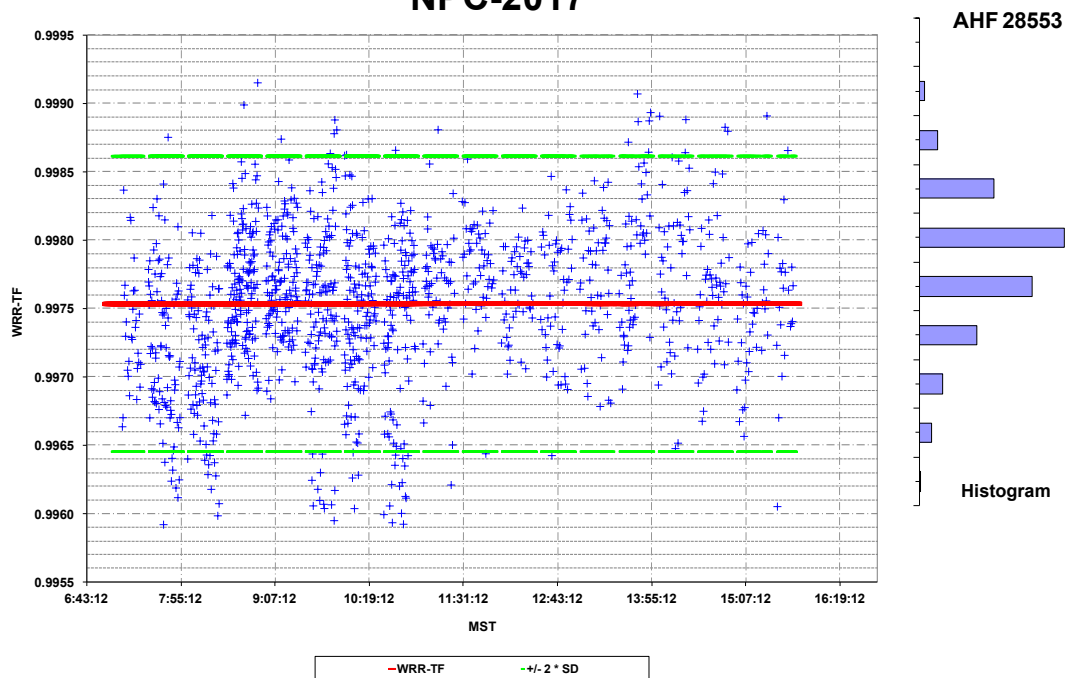


Figure 8. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF28553

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

AHF 28556

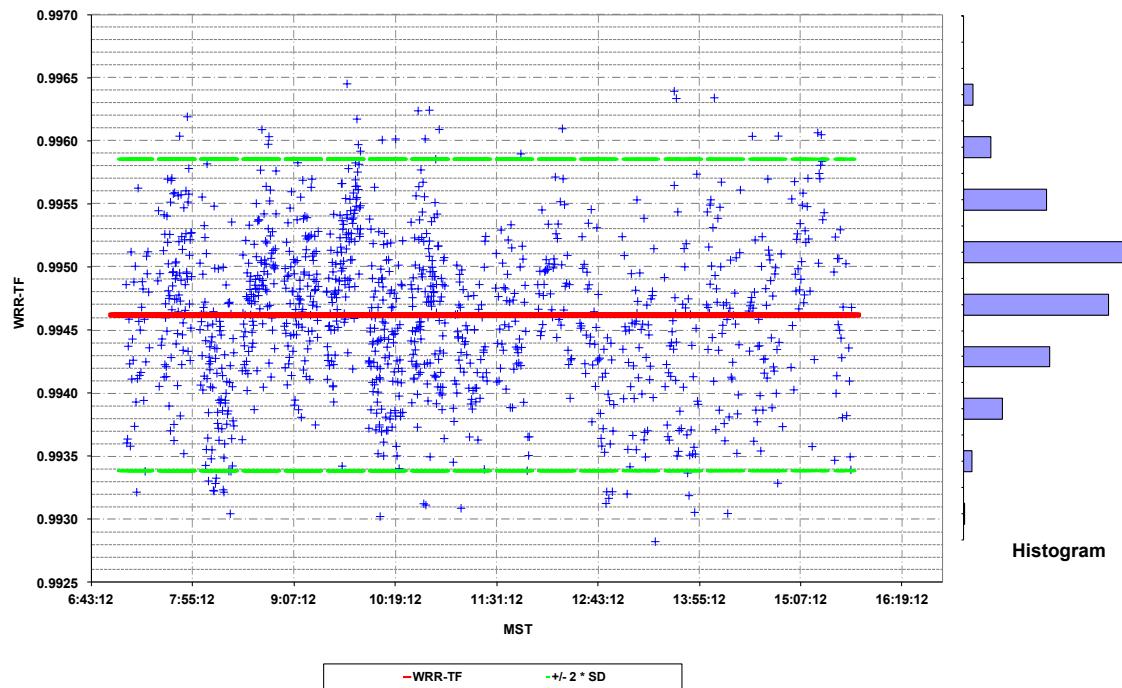


Figure 9. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 28556

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

AHF 28560

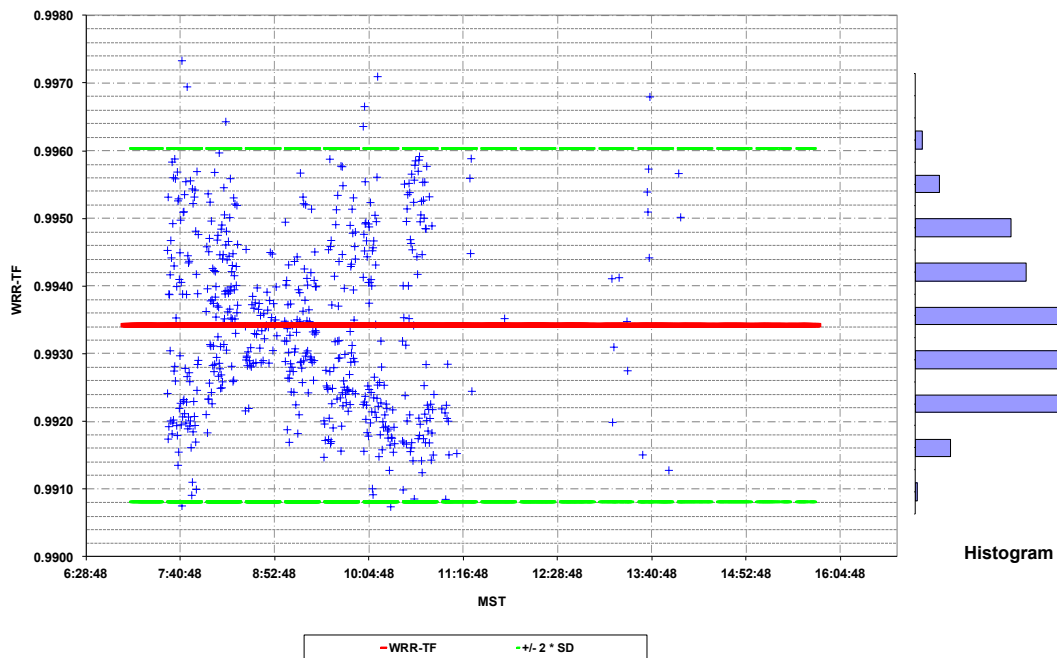


Figure 10. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 28560

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

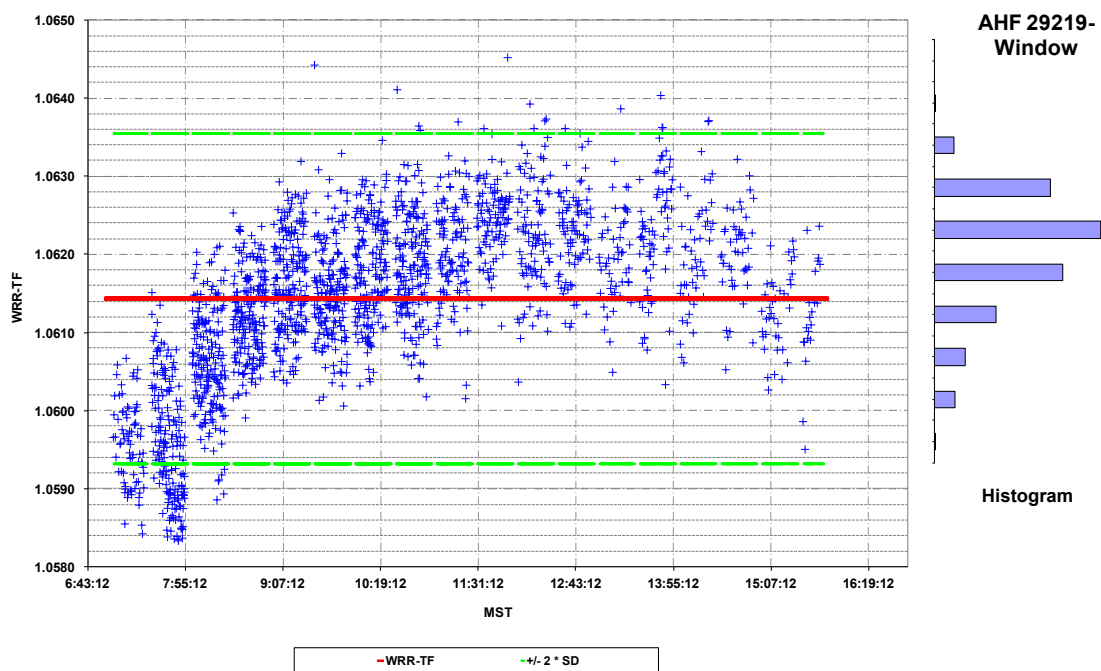


Figure 11. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 29219 - Windowed

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

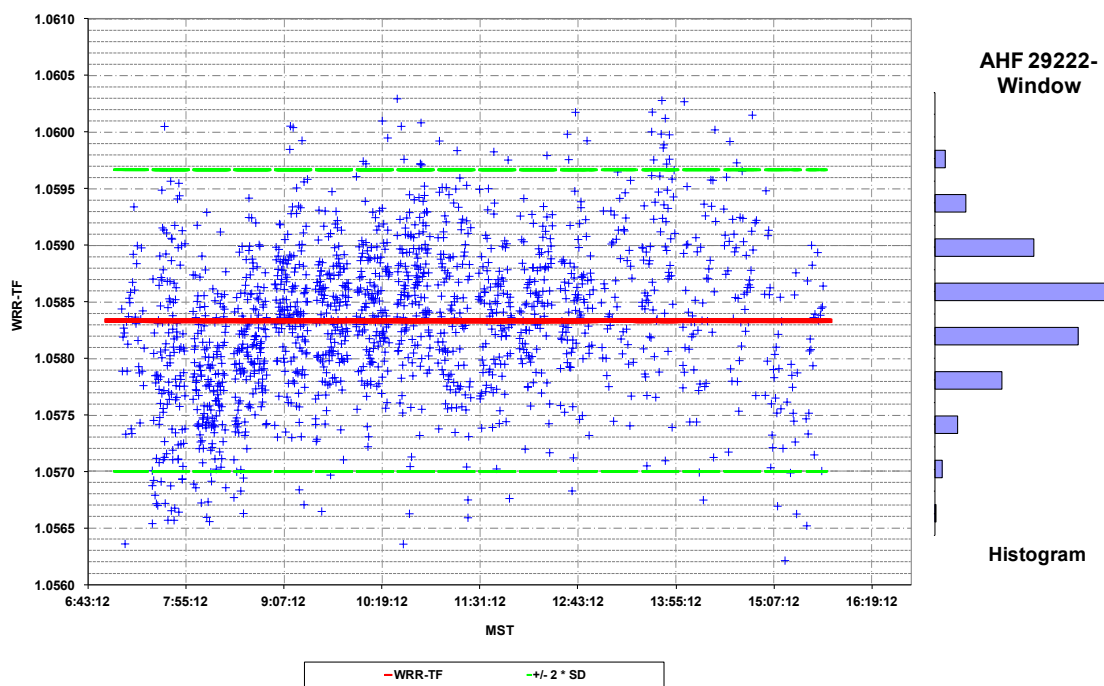


Figure 12. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 29222 – Windowed

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

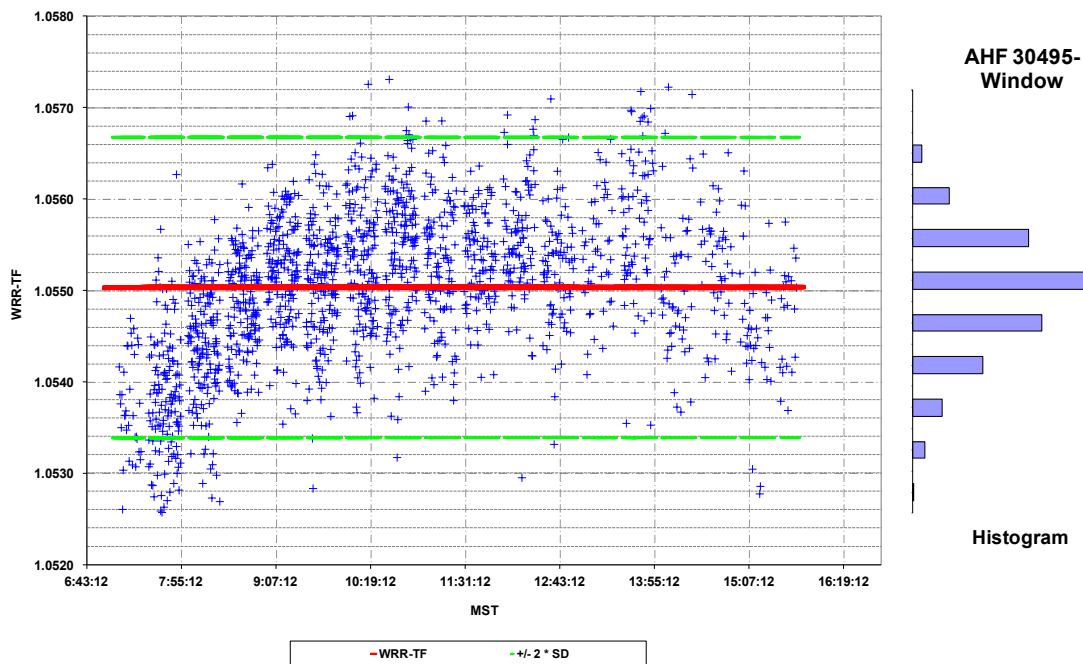


Figure 13. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF30495-Window

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

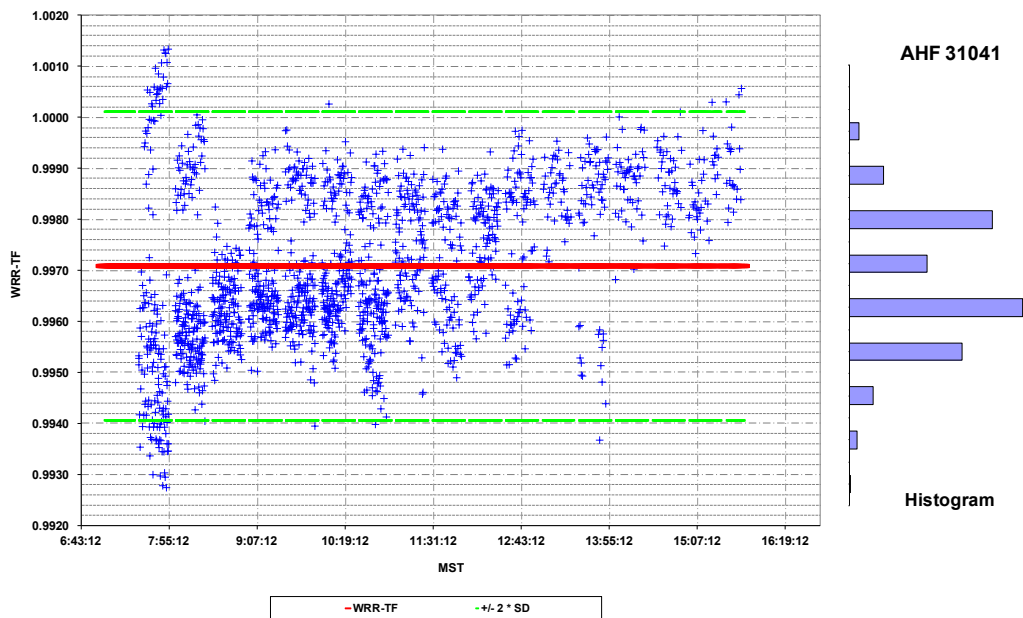


Figure 14. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31041

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

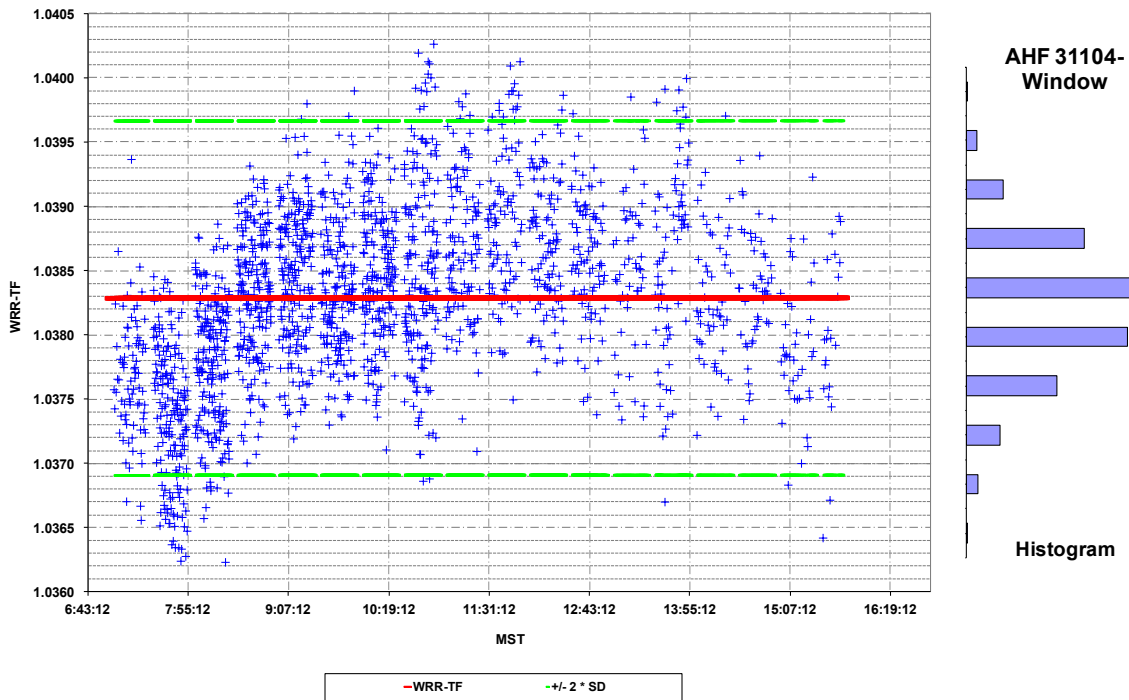


Figure 15. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31104-Window

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

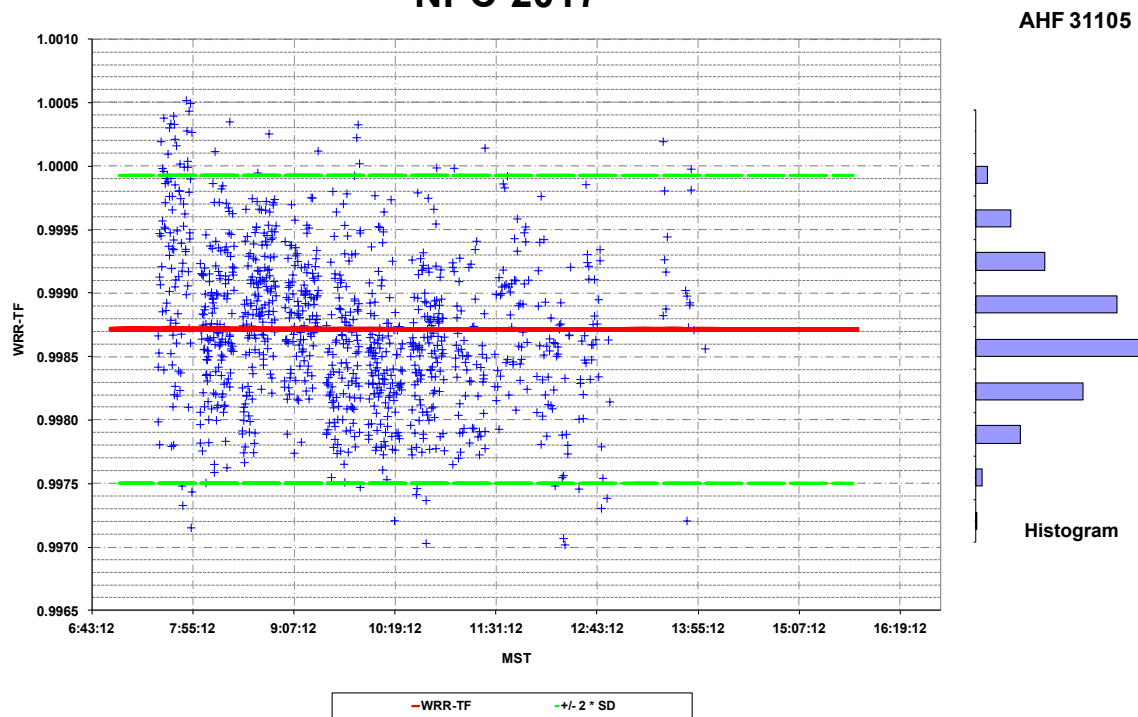


Figure 16. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31105

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

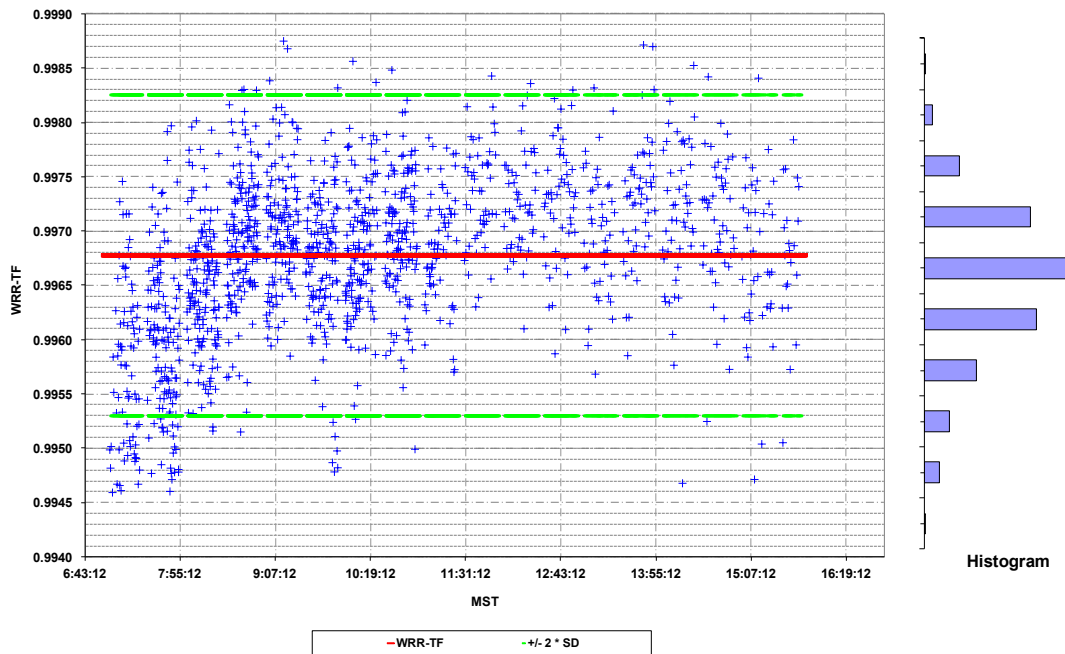


Figure 17. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31108

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

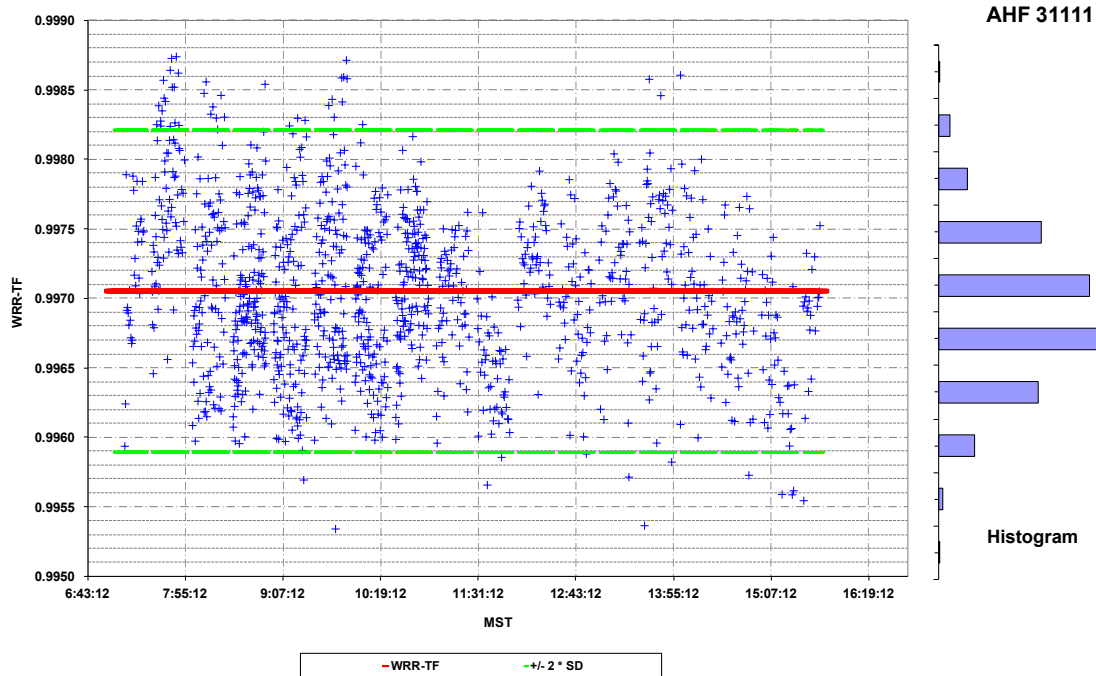


Figure 18. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31111

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

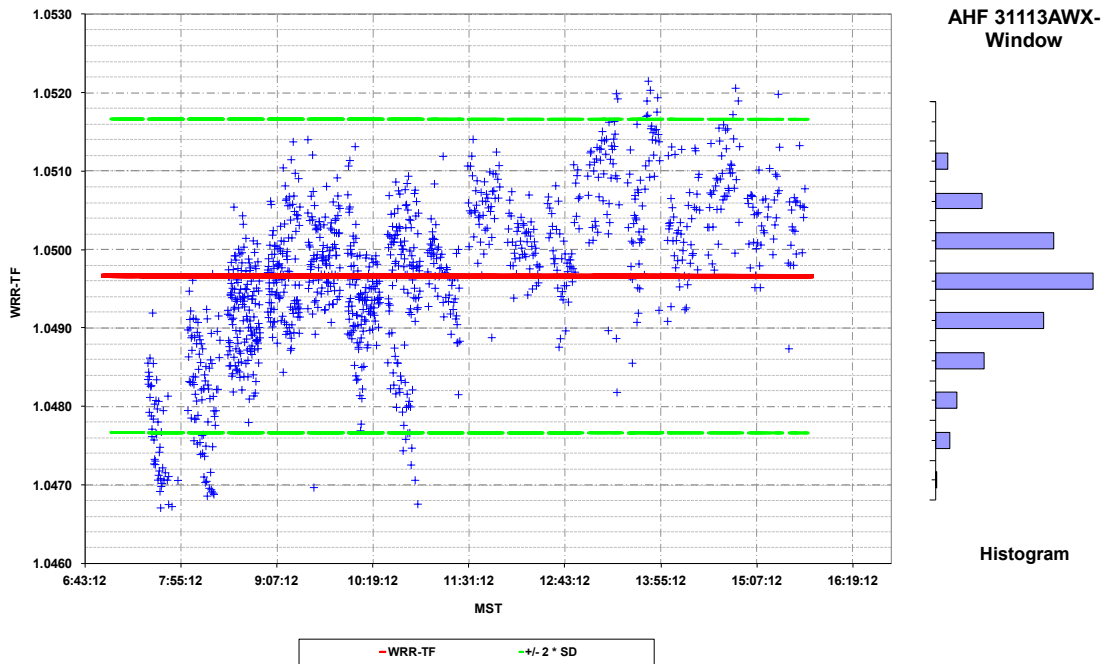


Figure 19. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31113AWX-Window

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

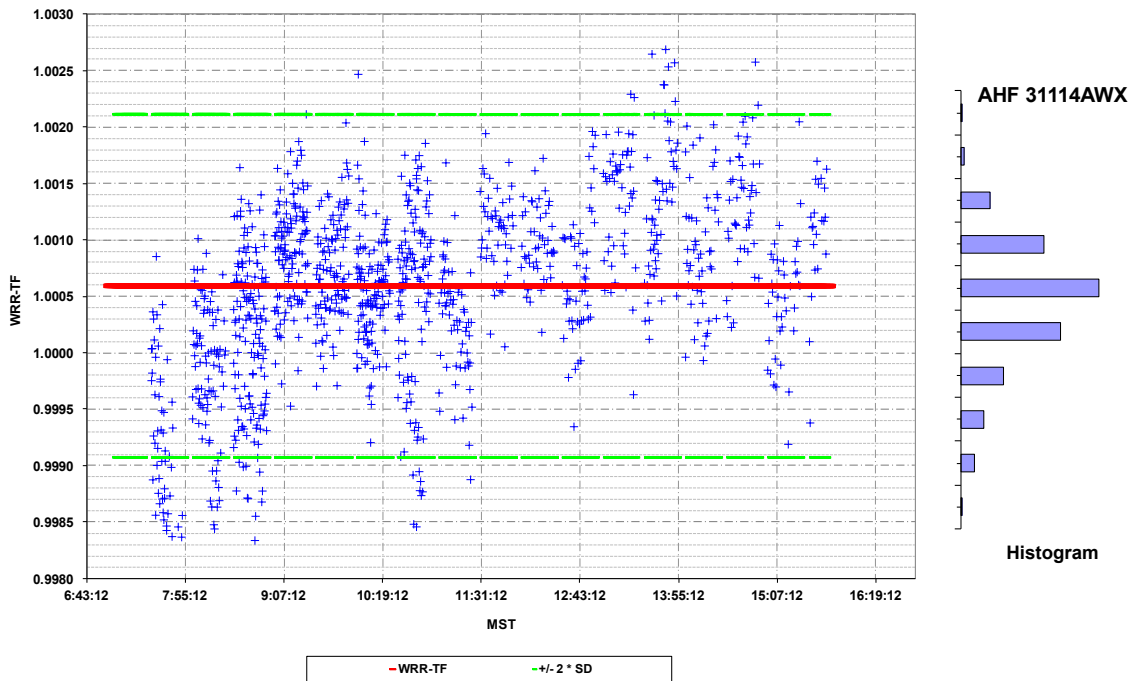


Figure 20. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31114AWX

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

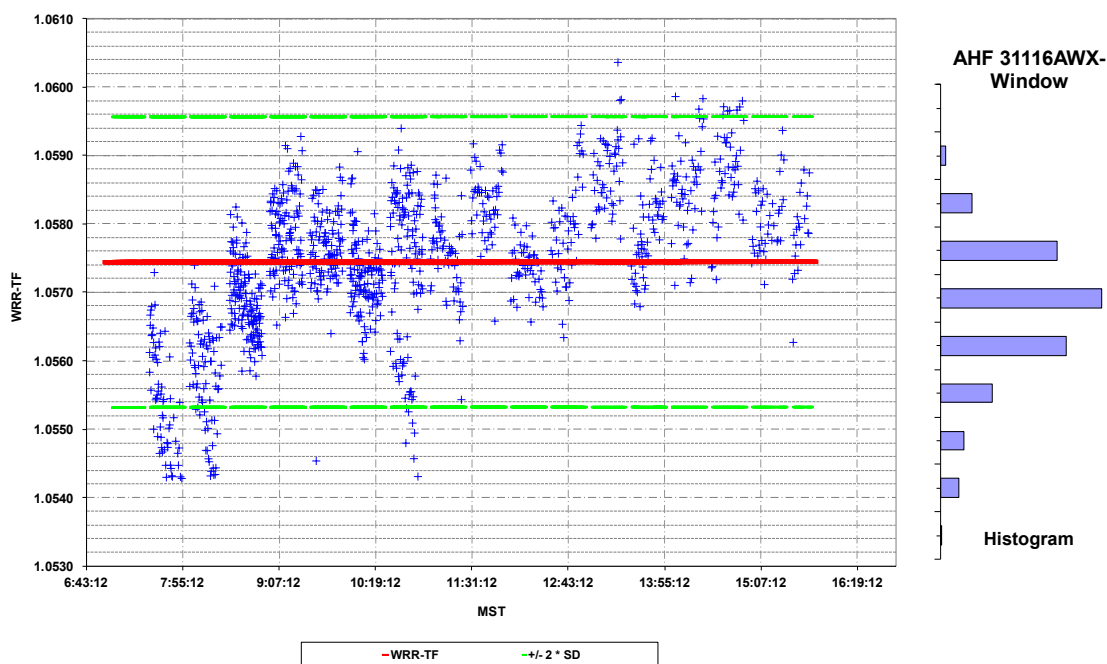


Figure 21. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 31116AWX-Window

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

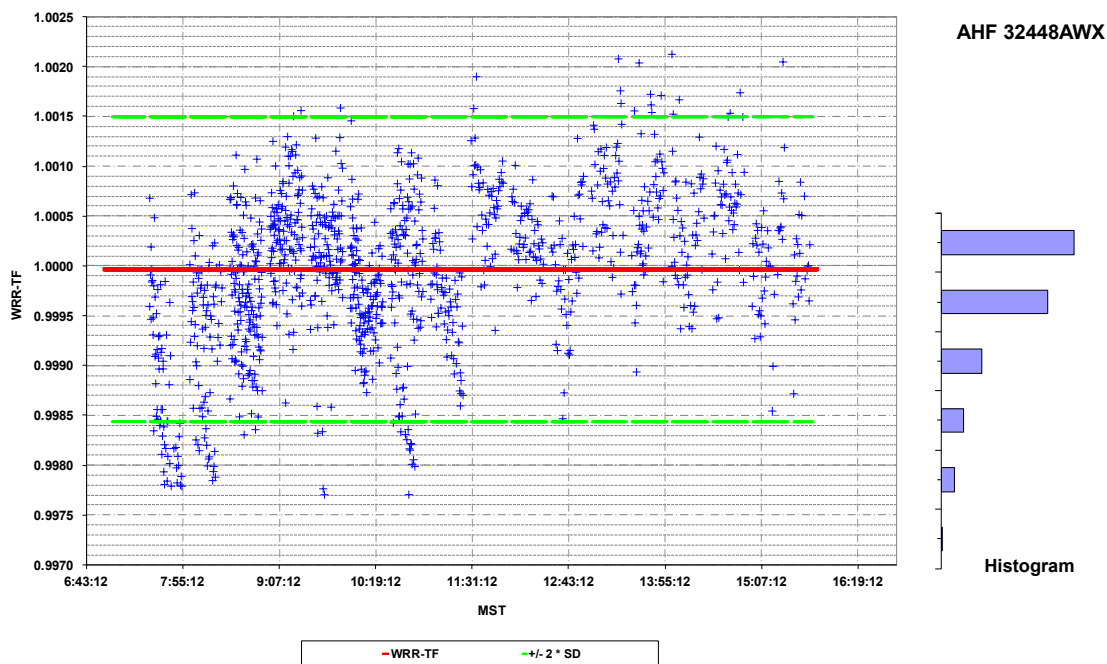


Figure 22. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 32448AWX

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

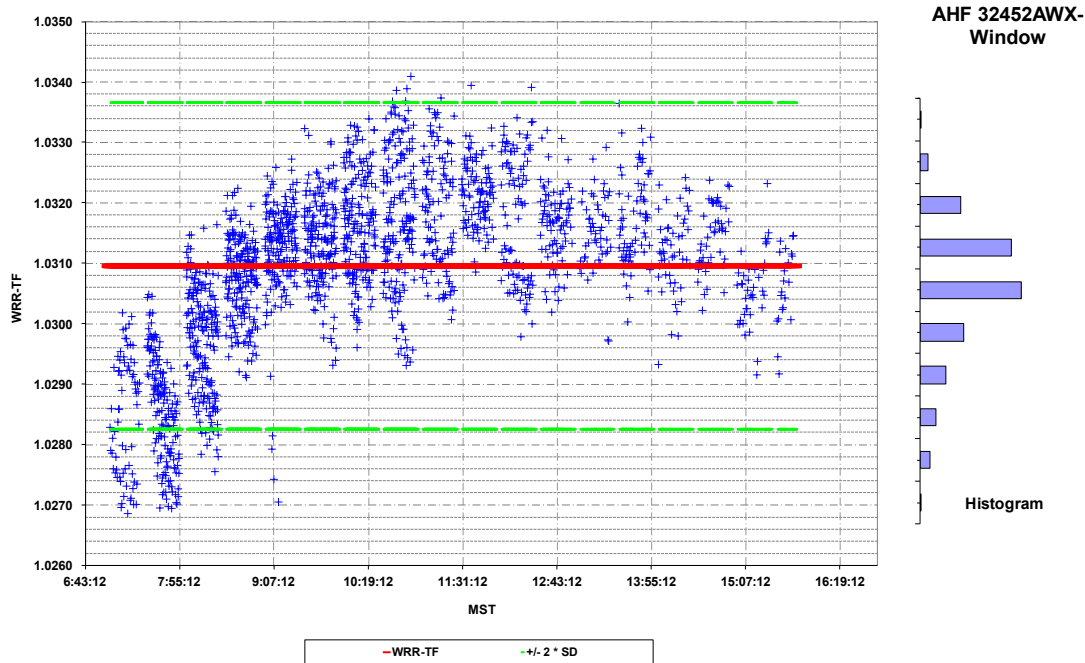


Figure 23. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 332452AWX – Windowed

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

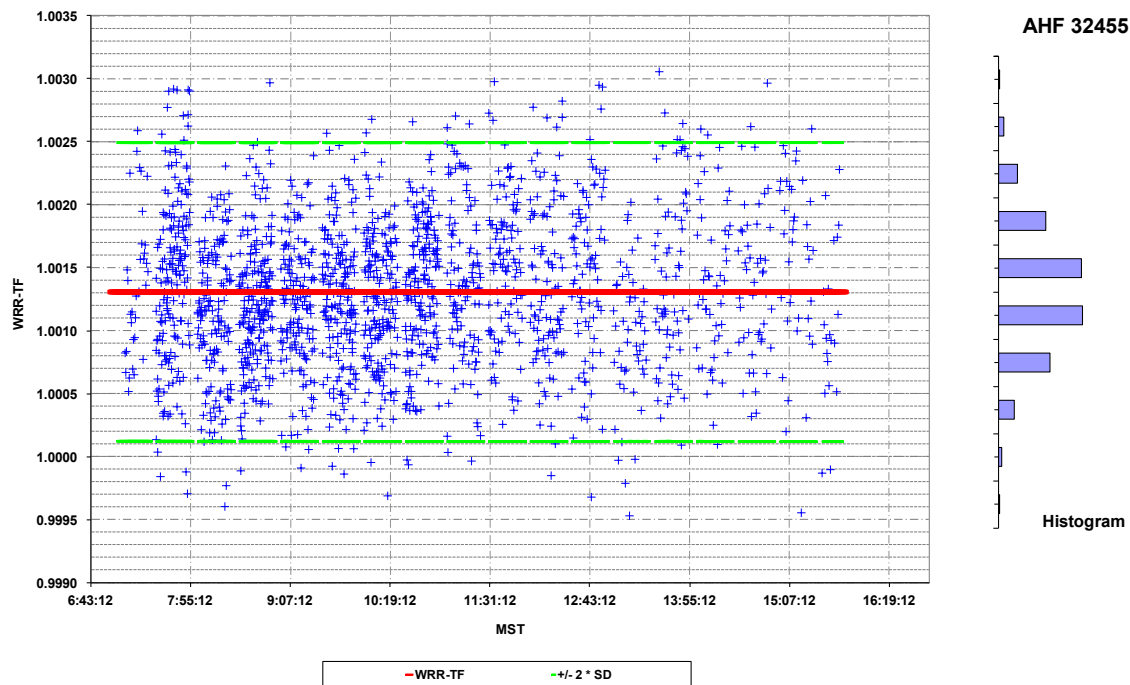


Figure 24. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 32455

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

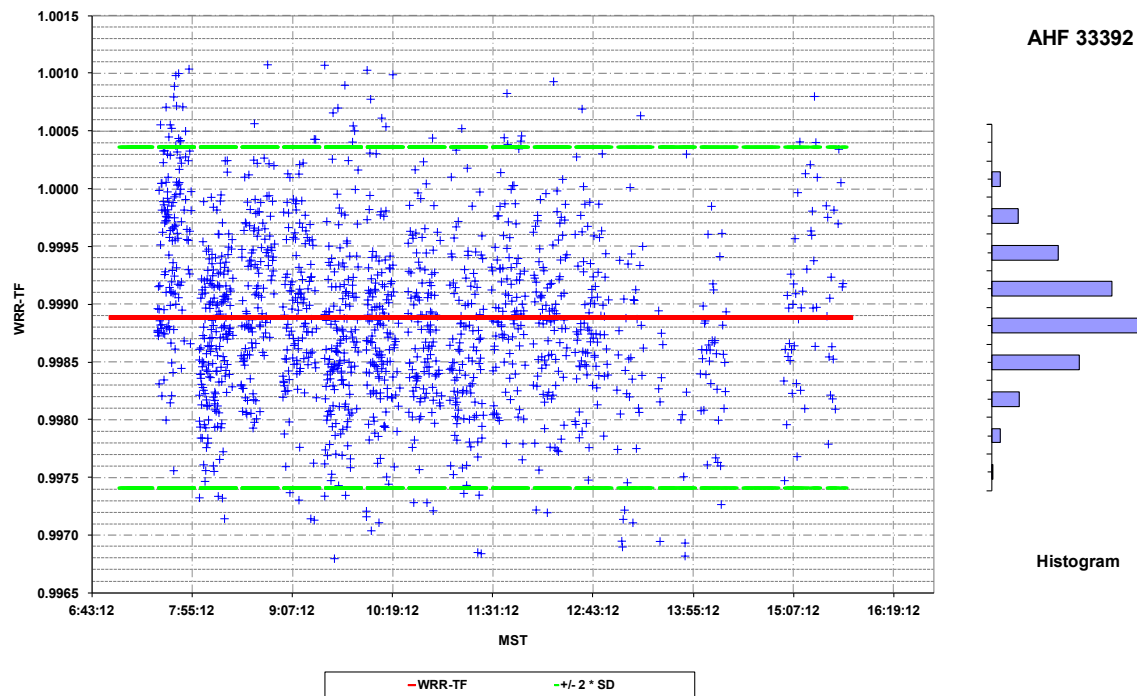


Figure 25. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 33392

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

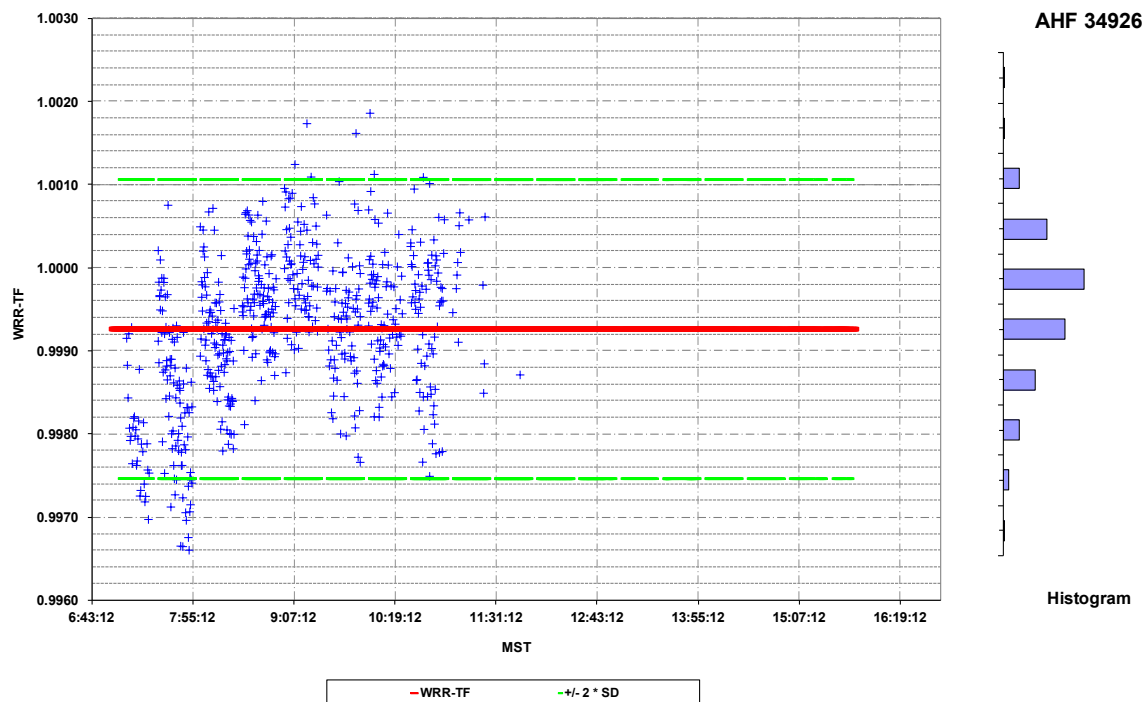


Figure 26. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 34926

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

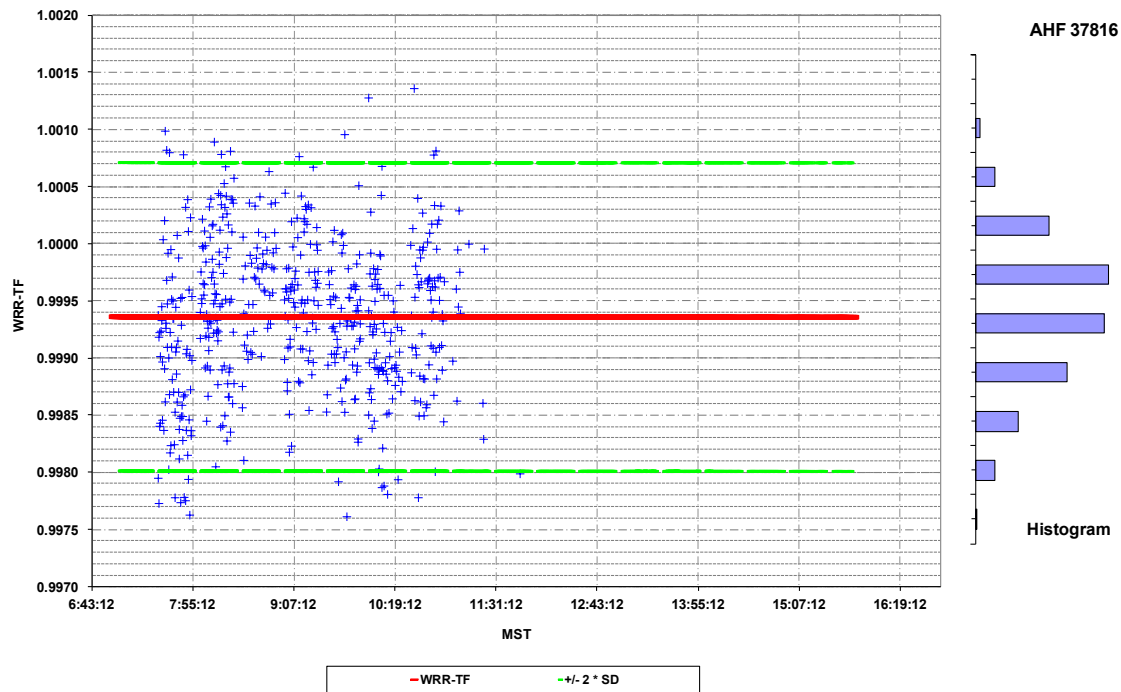


Figure 27. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for AHF 37816

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

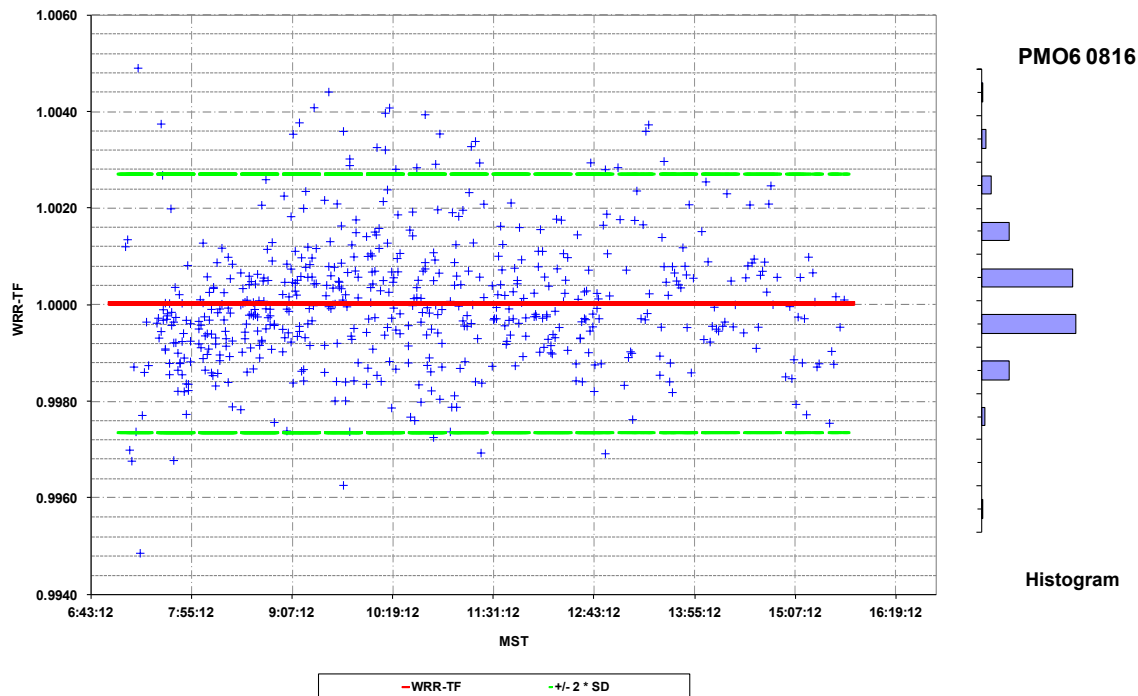


Figure 28. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for PMO6 0816

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

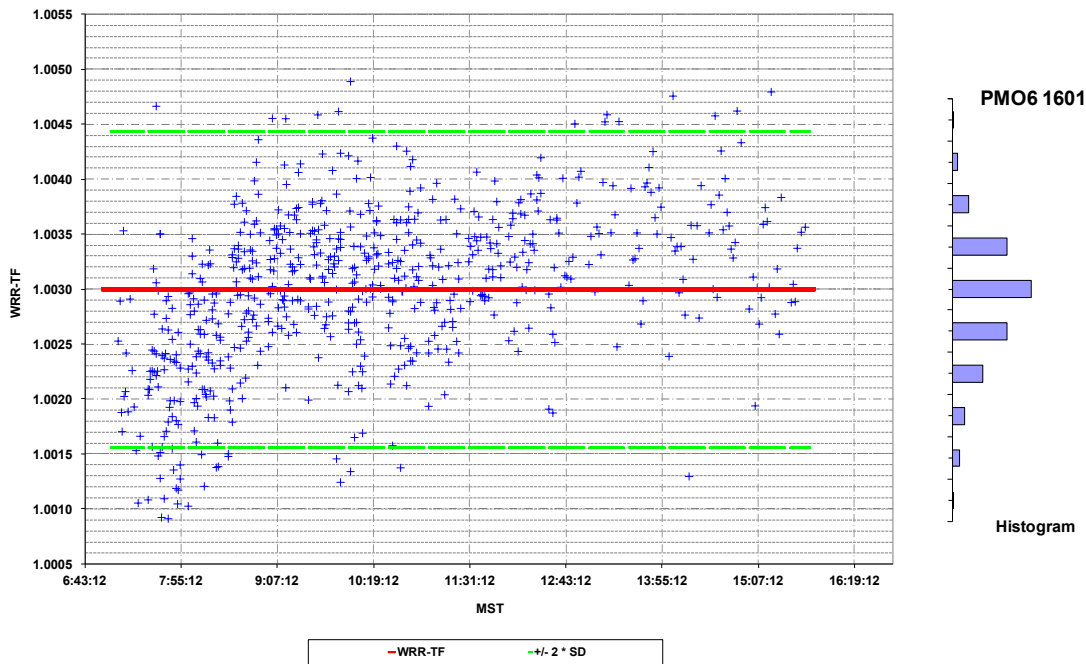


Figure 29. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for PMO6 1601

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

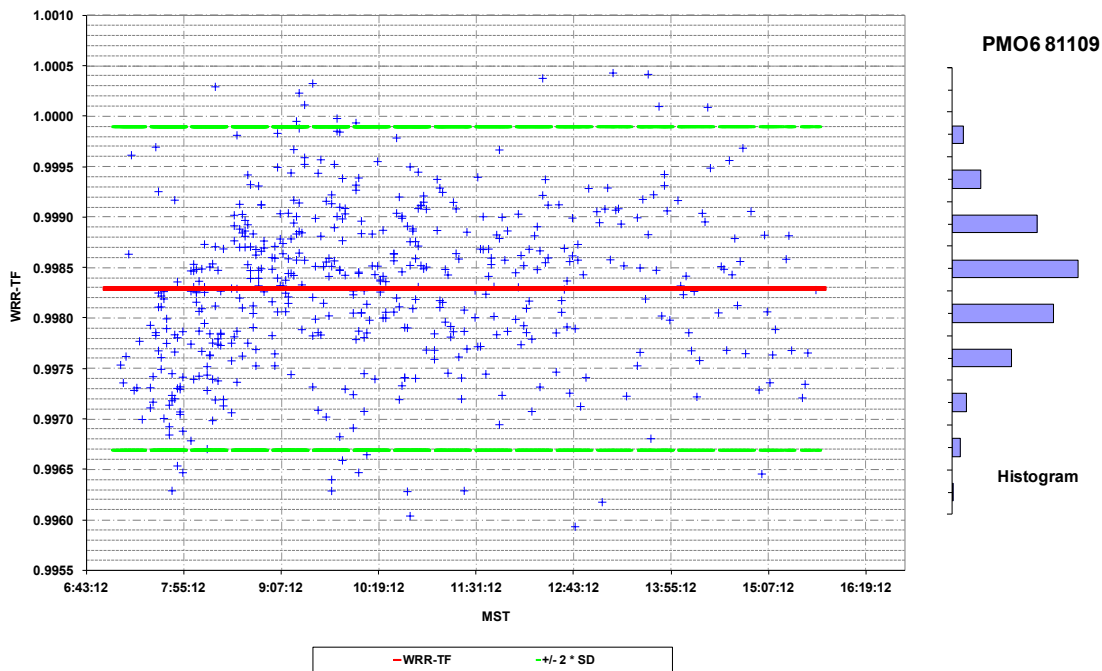


Figure 30. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for PMO6 81109

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

PMO6 911204

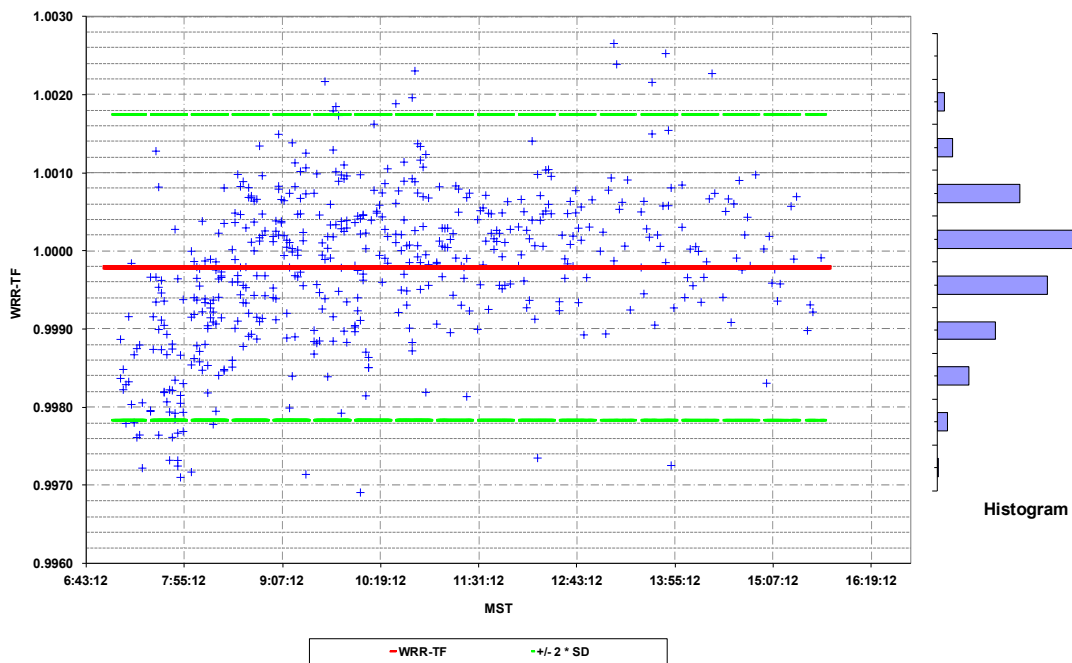


Figure 31. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for PMO6 911204

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

PMO6cc 0103

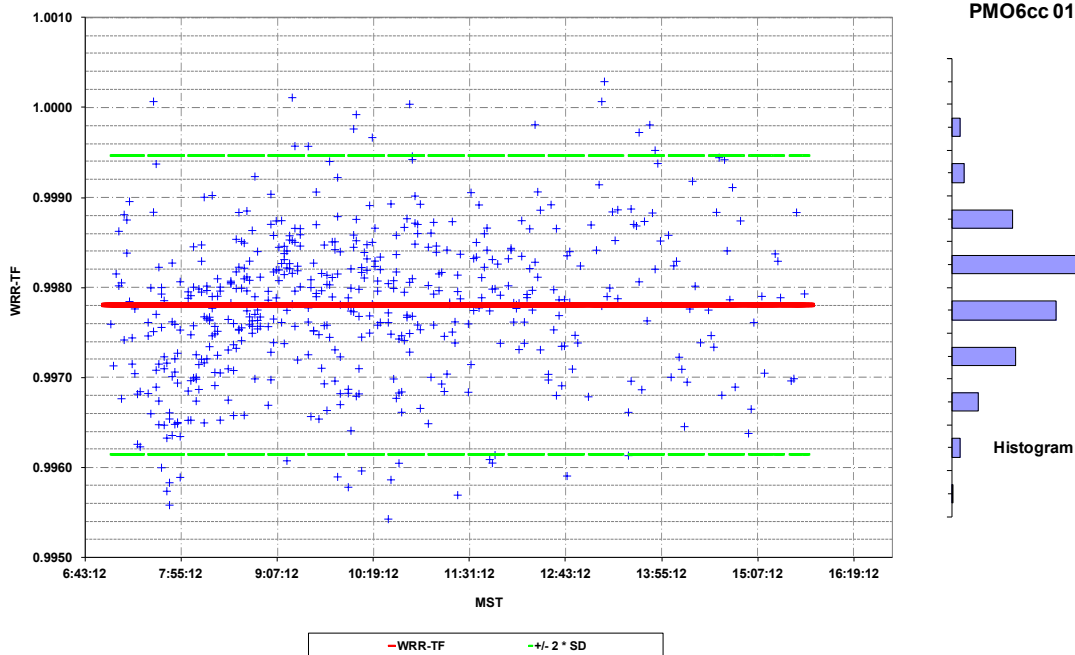


Figure 32. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for PMO6cc 0103

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

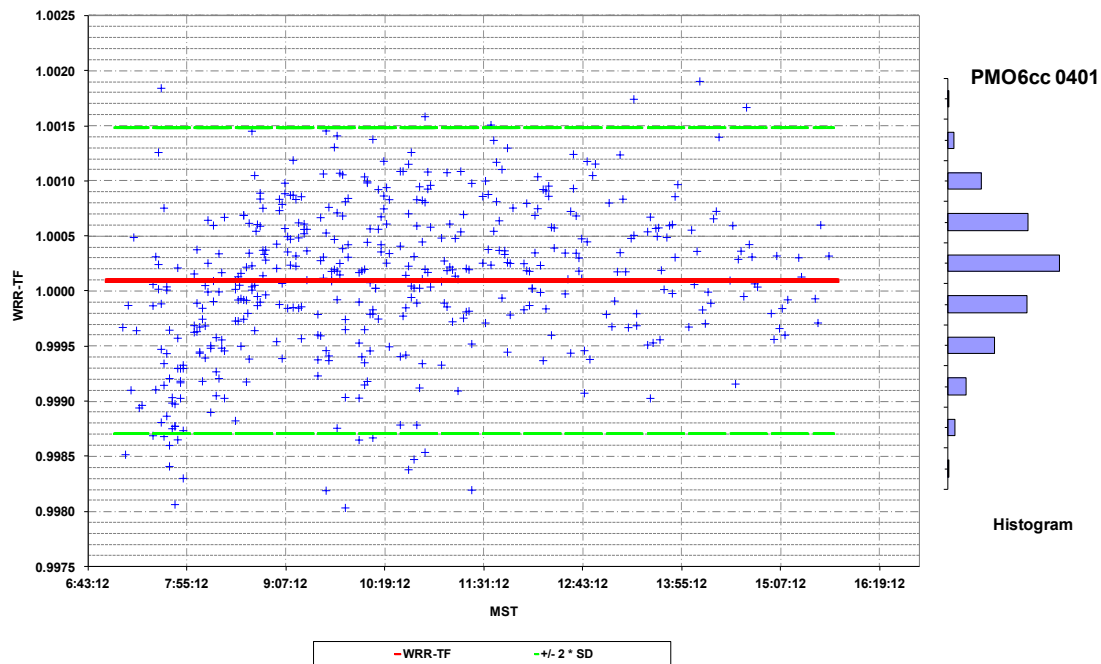


Figure 33. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for PMO6cc 0401

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

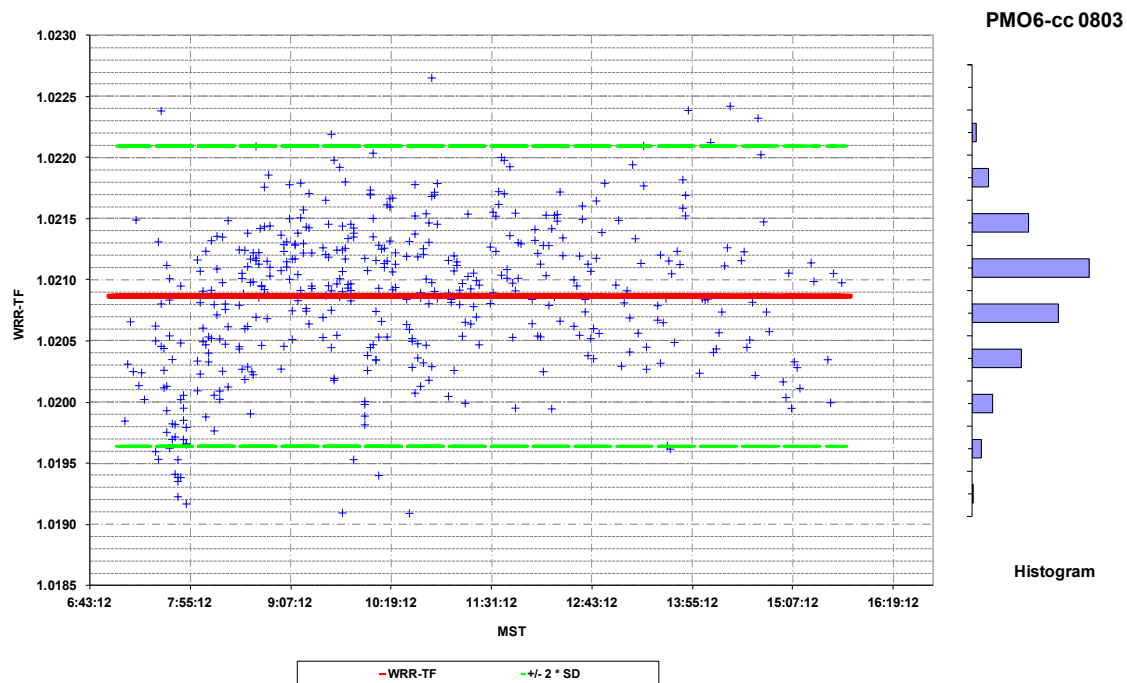


Figure 34. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for PMO6cc 0803

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

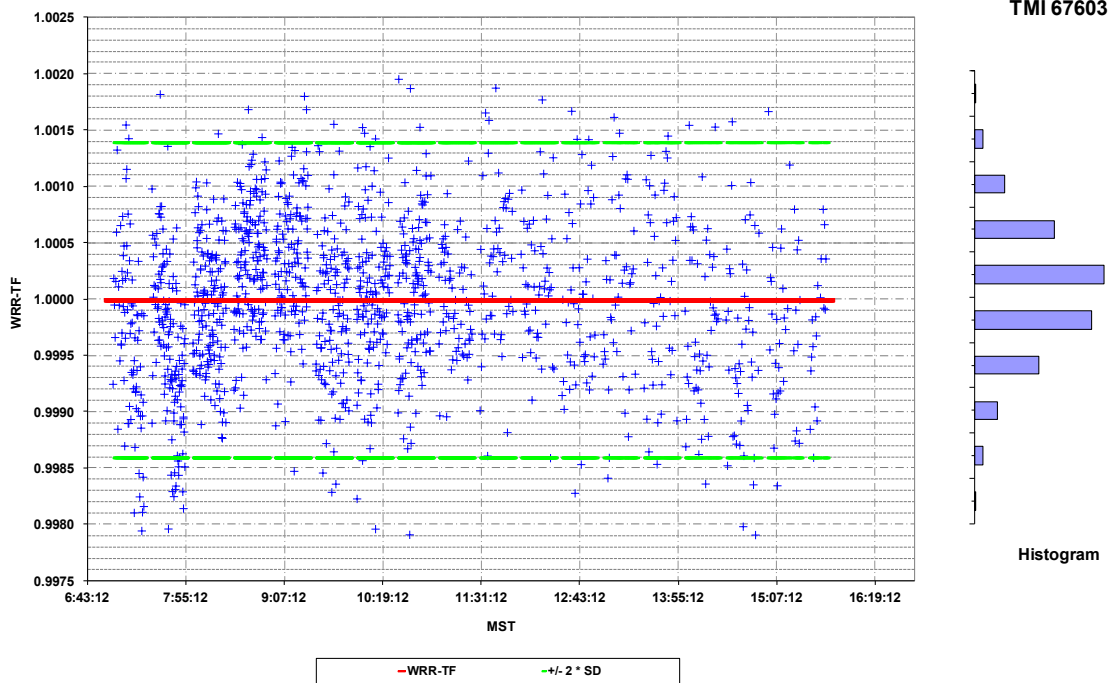


Figure 35. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for TMI 67603

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

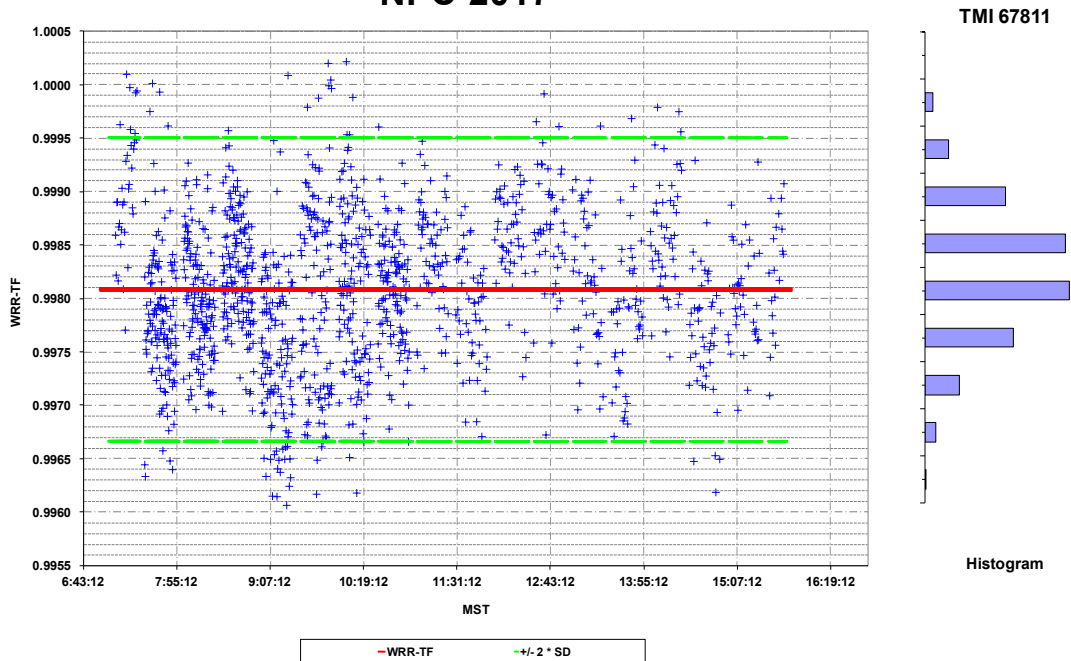


Figure 36. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for TMI 67811

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

TMI 68022

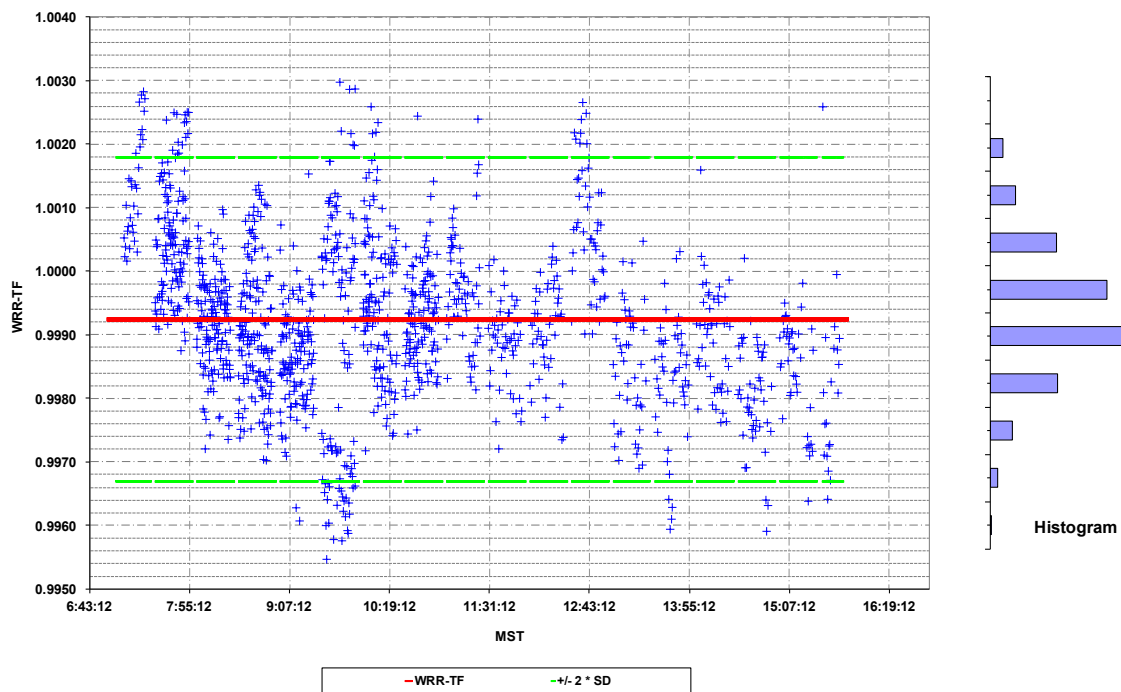


Figure 37. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for TMI 67811

WRR-Transfer Factor vs Mountain Standard Time NPC-2017

TMI 68835

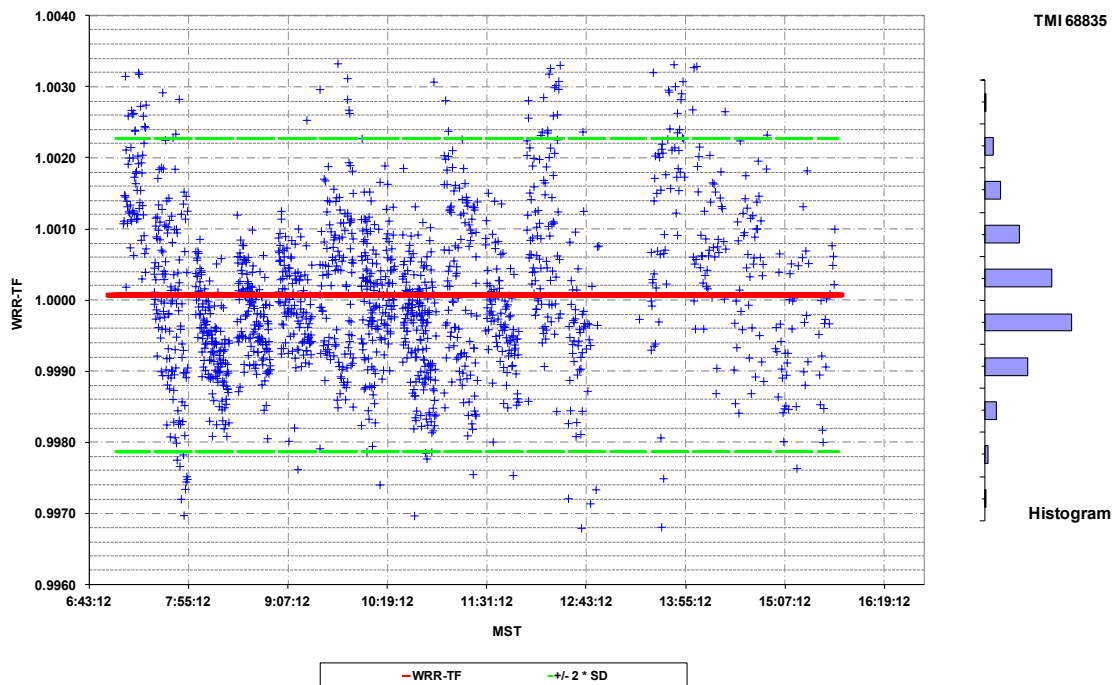


Figure 38. WRR-Transfer Factor vs. Mountain Standard Time NPC-2017 for TMI 68835

4.6 Recommendations

As a result of these comparisons, we suggest that participants observe the following measurement practices:

- For the purpose of pyrheliometer comparisons, such as NPC-2017, we recommend that the user apply only the manufacturer's calibration factor, not the WRR-TF or the new calibration factor, to report his or her absolute cavity radiometer's irradiance readings. This eliminates the possibility of compounding WRR factors from previous comparisons.
- For data collection in the field, the manufacturer's calibration factor should be used to calculate the cavity responsivity. Each irradiance reading should then be *multiplied* by the appropriate WRR-TF to provide homogeneity of solar radiation measurements that are traceable to the WRR. We recommend this approach to realize the benefits of participating in the NPC.

5 Ancillary Data

The environmental conditions (i.e., temperature, relative humidity, barometric pressure, wind speed, precipitable water vapor, and spectral data) were measured during the comparisons using the meteorological station at SRRL. Additional information, including data and graphical summaries, can be found at the Measurements and Instrumentation Data Center:

www.nrel.gov/midc/srrl_bms.

Time-series plots and other graphical presentations of these data collected during the pyrheliometer comparisons are presented in Appendix B.

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- WRC/PMOD (1996). *International Pyrheliometer Comparison, IPC VIII, 25 September – 13 October 1995, Results and Symposium*. Working Report No. 188. Davos Dorf, Switzerland: Swiss Meteorological Institute, Dorfstrasse 33, CH-7260; 115 pp.

Appendix A: List of Participants and Pyrheliometers

S/N	Operator 1	Operator 2	Affiliation
AHF 0000	Wim Zaaiman		European Commission Directorate General JRC
AHF 14915	Tom Kirk		EPPLEY LAB
AHF 14917	Emiel Hall	Jim Wendell	NOAA/ESRL/GMD
AHF 17142	Patrick Smith		Atlas Material Testing Technology, LLC
AHF 23734	Afshin Andreas	Mike Dooraghi	National Renewable Energy Laboratory
AHF 28553	Emiel Hall	Jim Wendell	NOAA/ESRL/GMD
AHF 28556	Patrick Smith		Atlas Material Testing Technology, LLC
AHF 28560	Erik Naranen		ISO-CAL North America, LLC
AHF 29219-Window	Andreas Afshin	Mike Doroaghi	National Renewable Energy Laboratory
AHF 29222-Window	Craig Webb		DOE Atmospheric Radiation Measurement (ARM) Program
AHF 30495-Window	Craig Webb		DOE Atmospheric Radiation Measurement (ARM) Program
AHF 31041	Fred Denn	Bryan Fabbri	Science Systems & Applications, Inc.
AHF 31104-Window	Afshin Andreas	Mike Dooraghi	National Renewable Energy Laboratory (Metrology, SePA, and Cell & Module Certification Group)
AHF 31105	Fred Denn	Bryan Fabbri	Science Systems & Applications, Inc.
AHF 31108	Bill Boyson	Charles Robinson	Sandia National Laboratories (Photovoltaic Systems Evaluation Lab)
AHF 31111	Sergio Cole	Emiel Hall	Universidade Federal de Santa Catarina
AHF 31113AWX-Window	Emiel Hall	Jim Wendell	NOAA/ESRL/GMD
AHF 31114AWX	Emiel Hall	Jim Wendell	NOAA/ESRL/GMD
AHF 31116AWX-Window	Emiel Hall	Jim Wendell	NOAA/ESRL/GMD
AHF 32448AWX	Emiel Hall	Jim Wendell	NOAA/ESRL/GMD
AHF 32452AWX-Window	Afshin Andreas	Mike Dooraghi	National Renewable Energy Laboratory

S/N	Operator 1	Operator 2	Affiliation
AHF 32455	Wolfgang Finsterle		PMOD/WRC
AHF 33392	Anthony Bucholtz	Elizabeth Reid	Naval Research Laboratory
AHF 34926AWX	Josh Peterson		SRML University of Oregon
AHF 37816	Erik Naranen		ISO-CAL North America, LLC
PMO6 0816	Akihito Akiyama	William Beuttell	EKO Instruments USA, Inc
PMO6 1601	Ajay Singh	Matthew Perry	Campbell Scientific Inc.
PMO6 81109	Wim Zaaiman		European Commission Directorate General JRC
PMO6 911204	Wim Zaaiman		European Commission Directorate General JRC
Pmo6cc 0103	Victor Cassella	Joop Mes	Kipp & Zonen USA, Inc
PMO6cc 0401	Wolfgang Finsterle		PMOD/WRC
PMO6-cc 0803	Wolfgang Finsterle		PMOD/WRC
TMI 67603	Bill Boyson	Charles Robinson	Sandia National Laboratories (Photovoltaic Systems Evaluation Lab)
TMI 67811	Tim Moss		Sandia National Labs (Concentrating Solar Energy)
TMI 68022	Tim Moss		Sandia National Labs (Concentrating Solar Energy)
TMI 68835	Wim Zaaiman		European Commission Directorate General JRC

Appendix B: Ancillary Data Summaries

The measurement performance of an absolute cavity can be affected by several environmental parameters. Potentially relevant meteorological data collected during the NPC are presented in this appendix. The BMS has been in continuous operation at the SRRL since 1981. BMS data are recorded as 1-minute averages of 3-second samples for each instrument. (Additional information about SRRL and the BMS can be found at the Measurement and Instrumentation Data Center: http://www.nrel.gov/midc/srrl_bms/).

Time-series plots and other graphical presentations of these data acquired during the NPC-2017 measurements are presented here.

